

ISARS – June 2010 – Paris.

First Results of two Optical and Millimeter-wave Scintillometer systems during LITFASS2009

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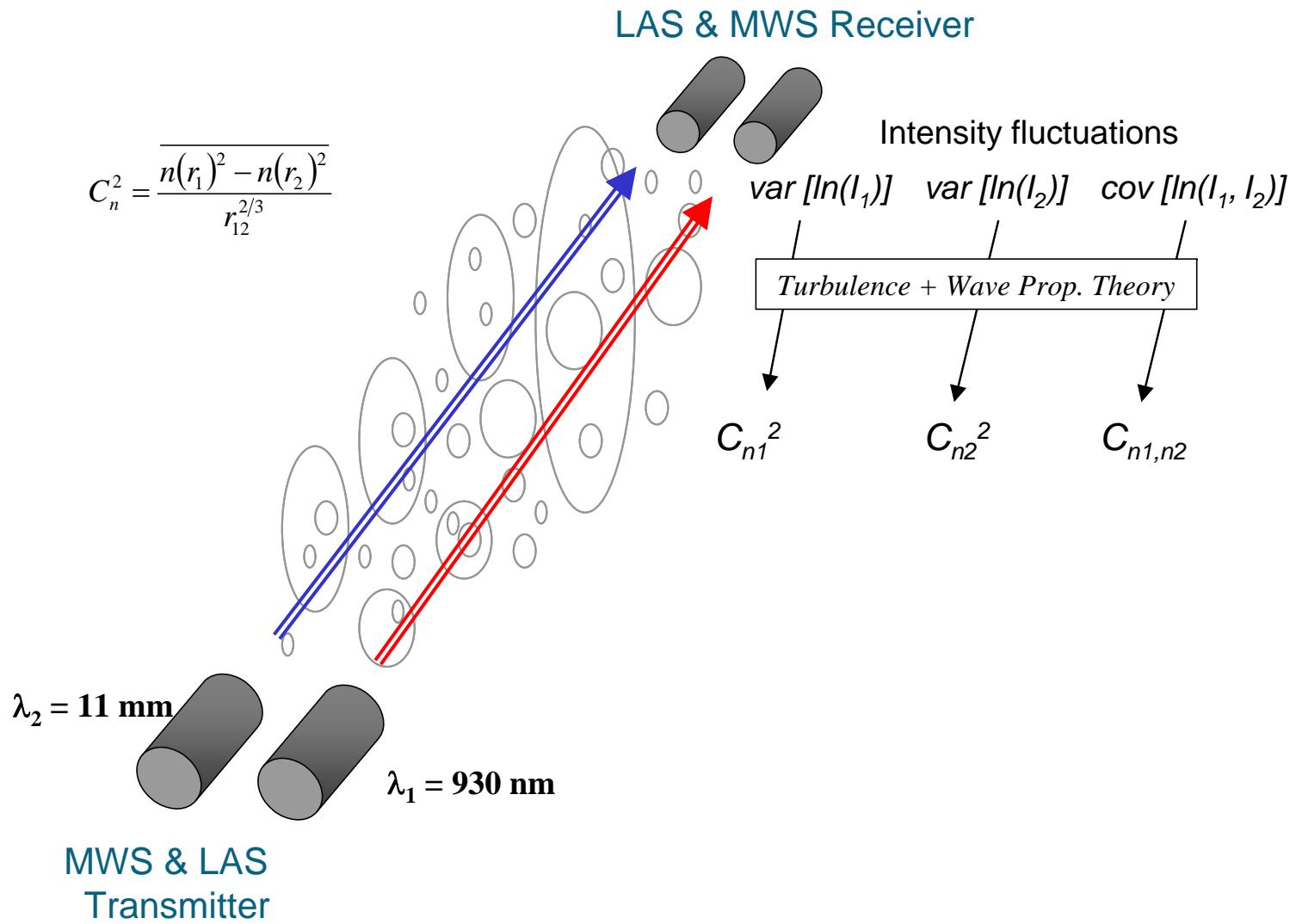
Outline Talk:



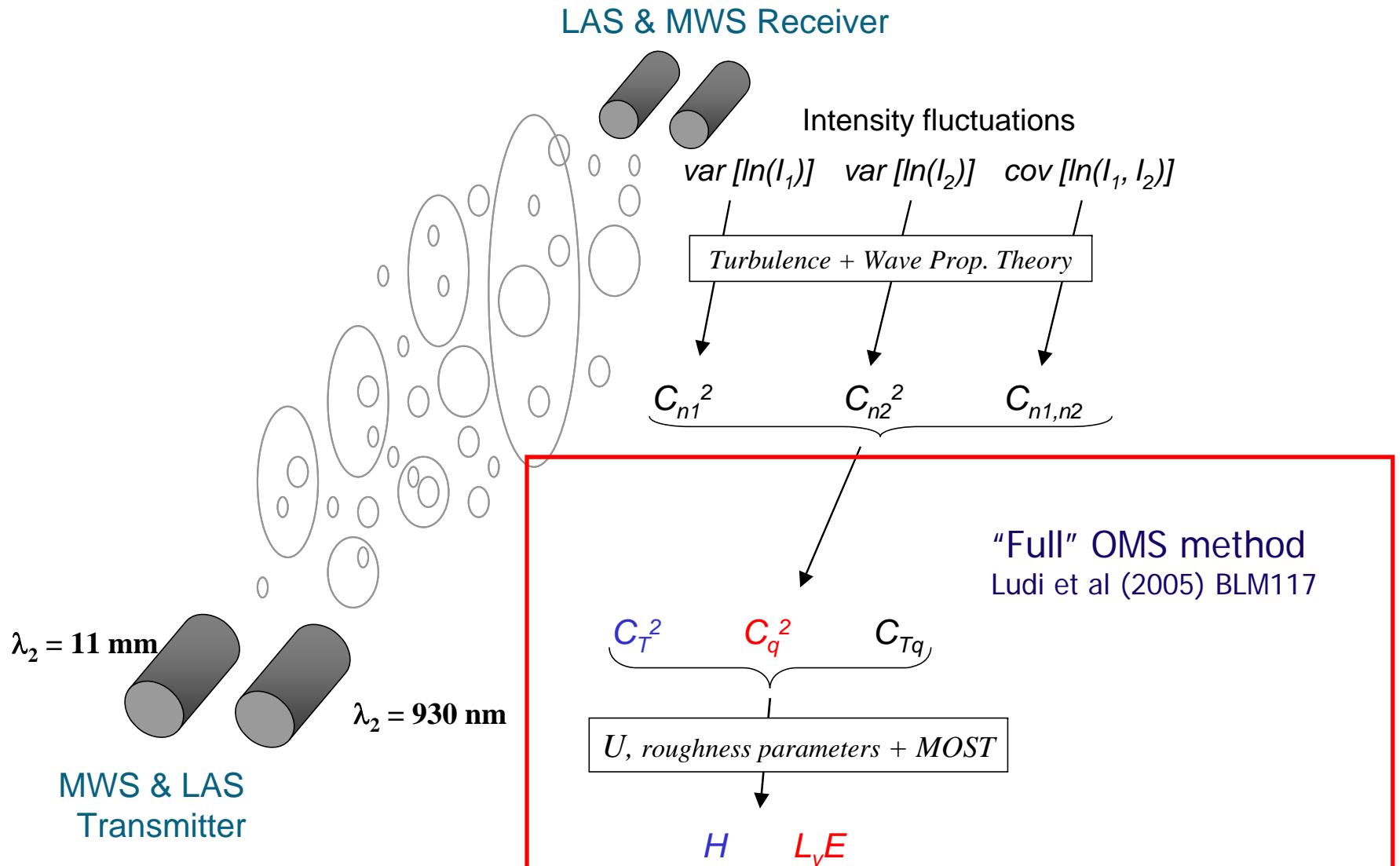
- Describe an Optical-Millimetre wave Scintillometer (OMS) System that measures $H + L_\nu E$ directly at km-scale.
- First results of LITFASS2009 experiment – Cross-wind.



OMS System Description



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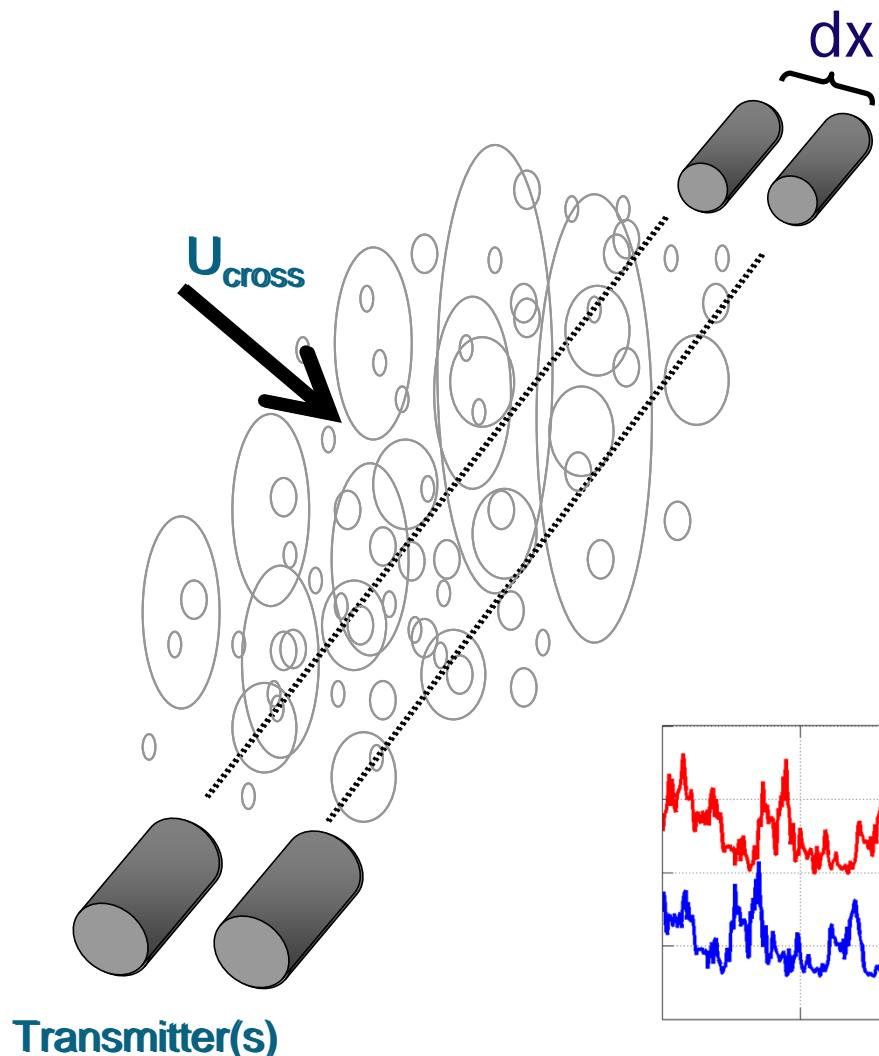


LITFASS2009 Experiment – DWD Meteorological Observatory Lindenberg



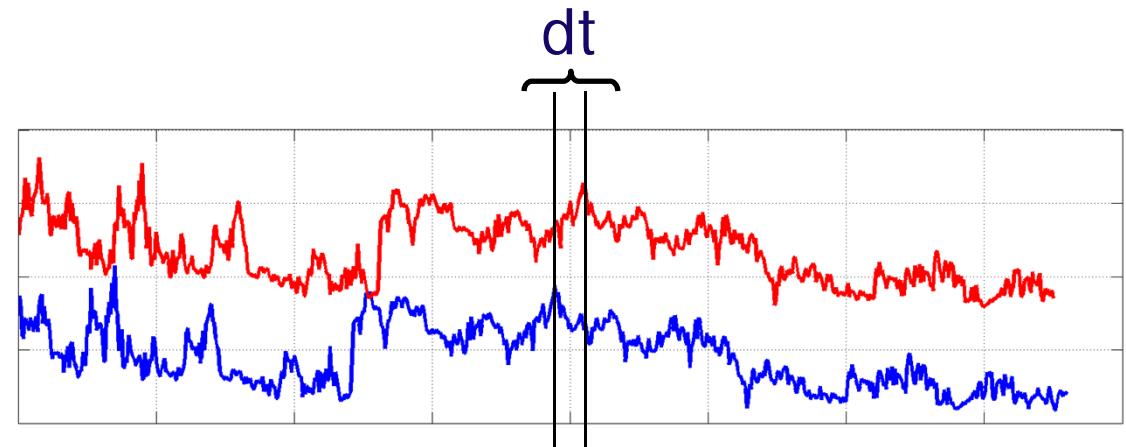
- CEH94 MV = $\lambda = 740\text{nm}$, $D = 0.15\text{m}$ = receiver frequency
- MOL94 MV = $\lambda = 740\text{nm}$, $D = 0.15\text{m}$ = receiver frequency
- Wageningen LWS = $\lambda = 740\text{nm}$, $D = 0.15\text{m}$ = concave mirror to focus beam onto detector
- Path length $W = 10\text{m}$ Raw data @ 500Hz
- and now for something completely different: CROSS-WIND

Cross-wind – Dual Aperture Approach



Scintec BLS-series

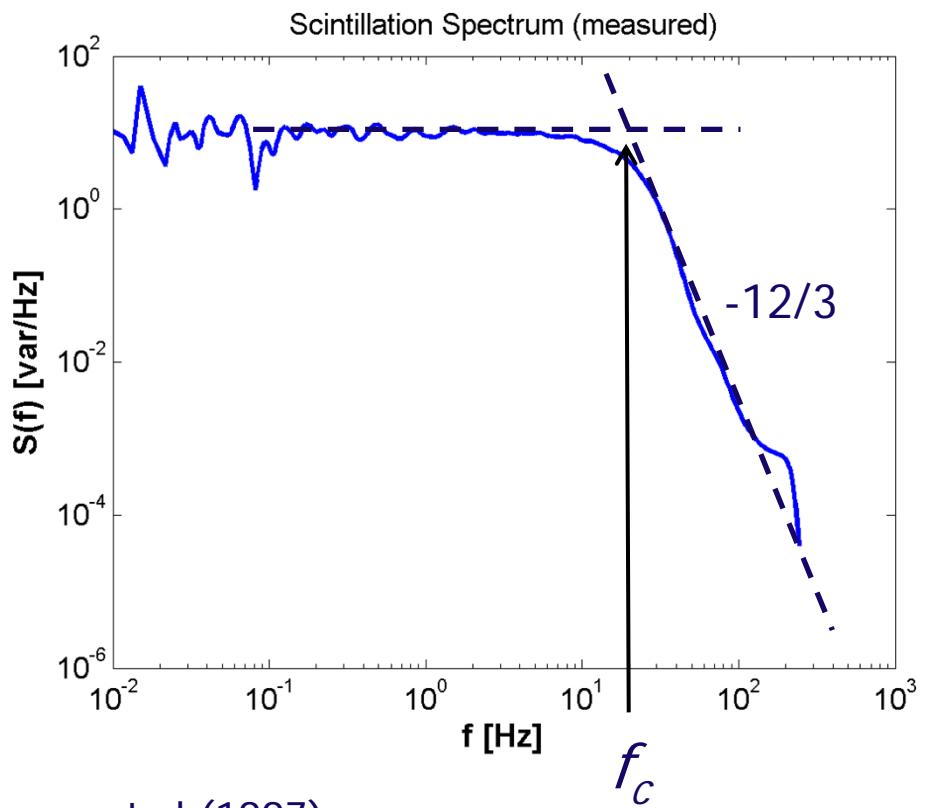
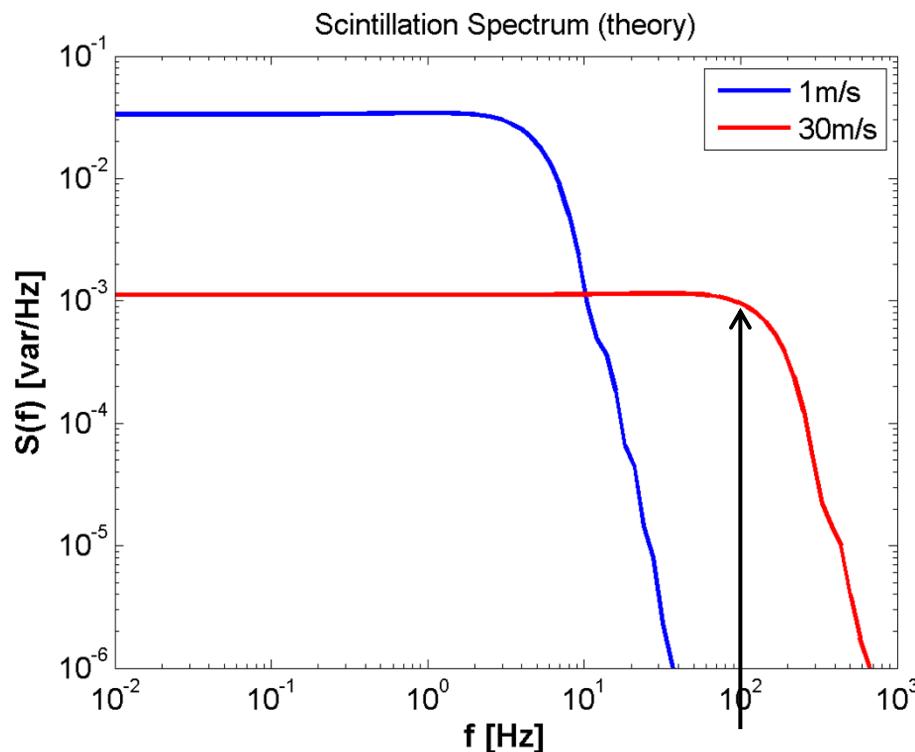
$$U_{cross} = dx/dt$$



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Cross-wind – Single Aperture Approach1 – Corner-Frequency

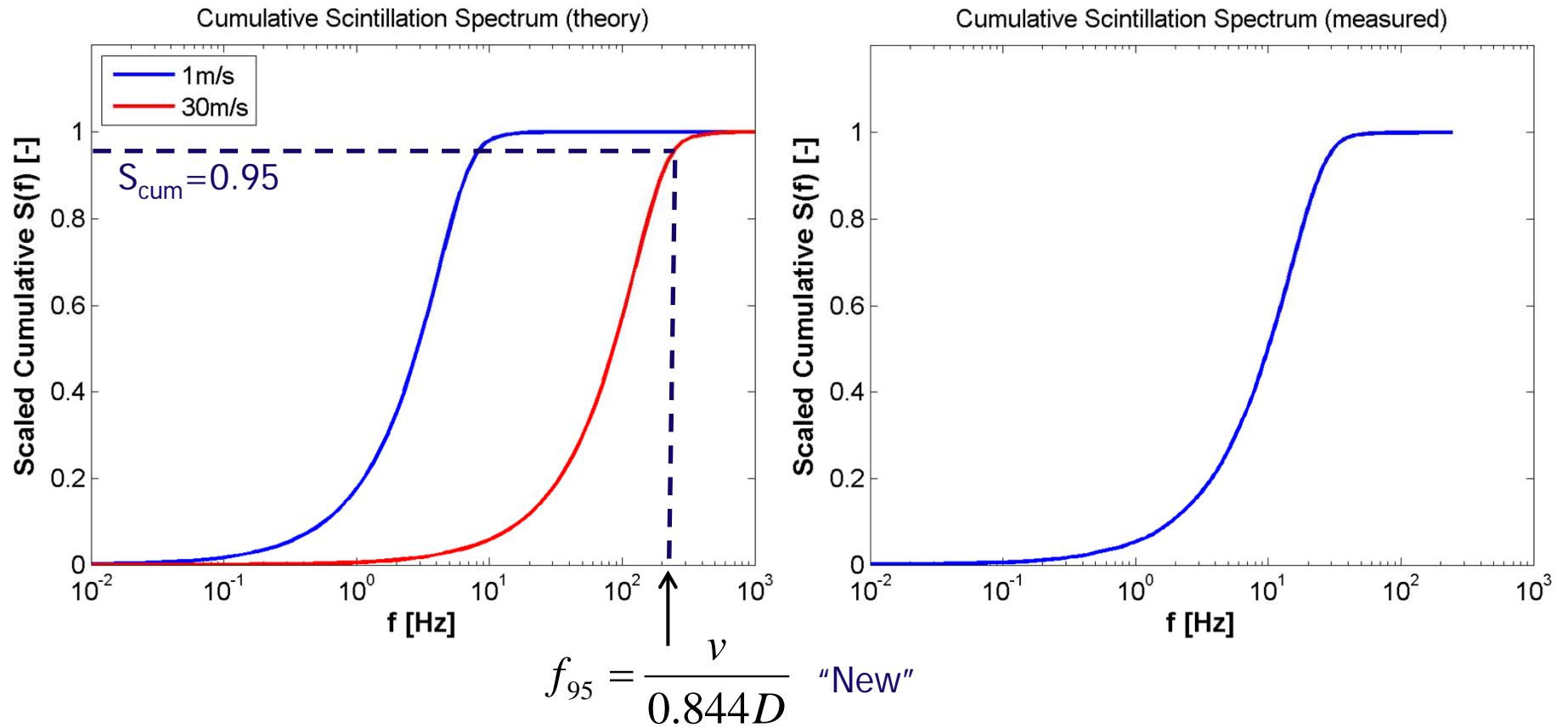
Scintillometer Spectrum: $S(f) = 4\pi^2 K^2 \int_0^L \int_{2\pi f/v}^{\infty} k \phi_n \sin^2 \left(\frac{k^2 x(L-x)}{2KL} \right) [(kv)^2 - (2\pi f)^2]^{-1/2} dk dx$
Clifford (1971)



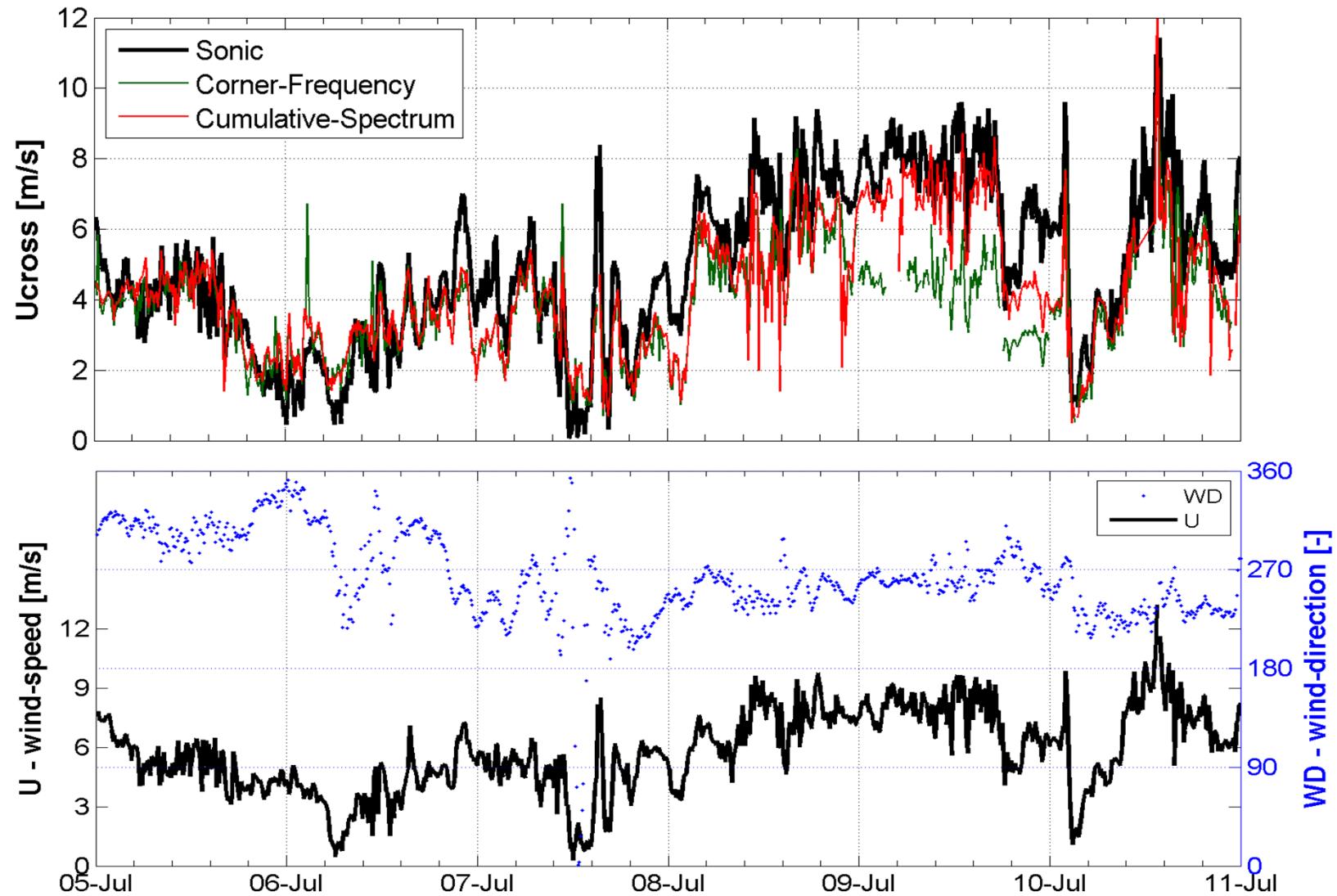
$$f_c = \frac{v}{1.25D}$$
 Nieveen et al (1997)



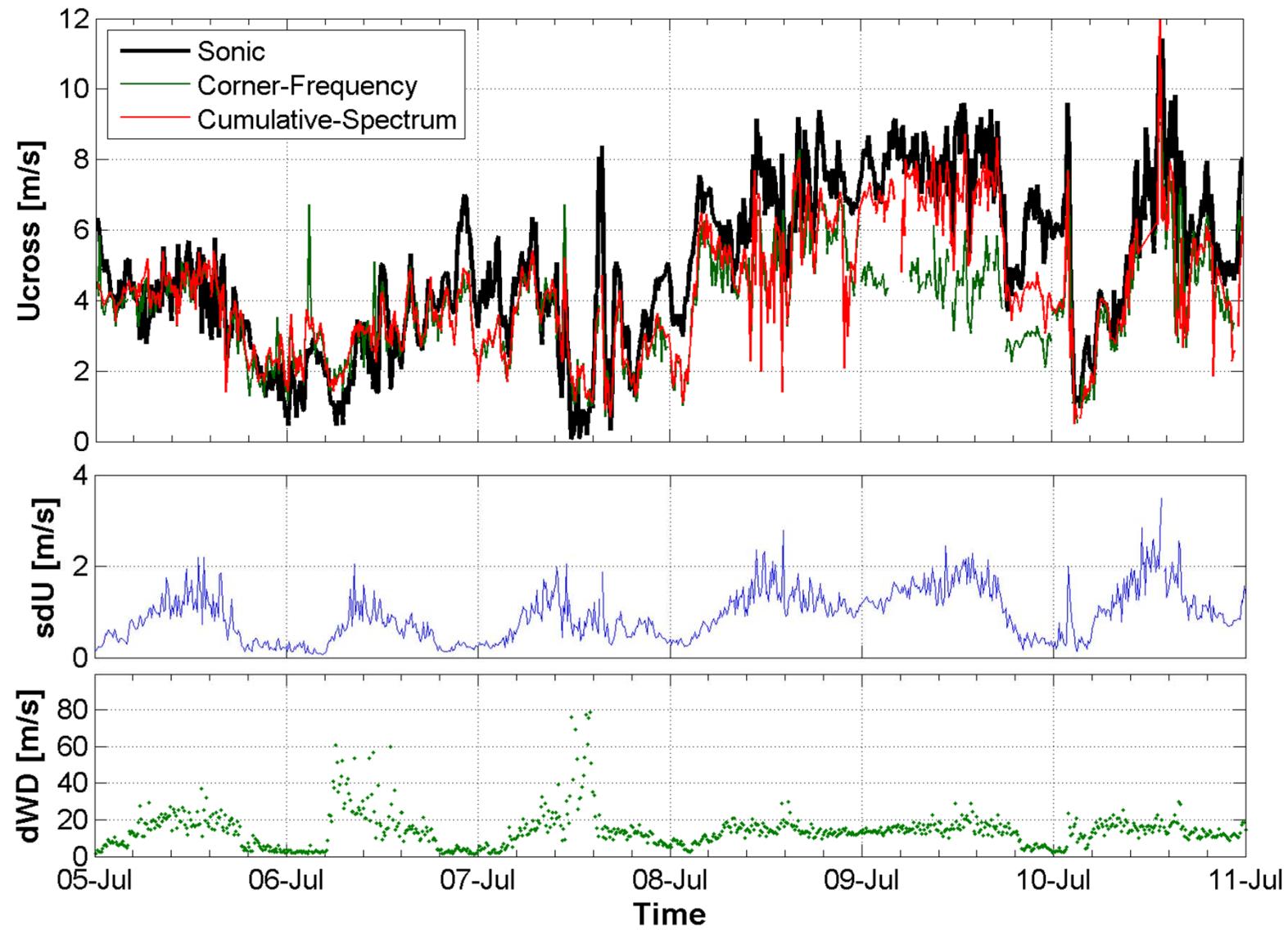
Cross-wind – Single Aperture Approach2 – Cumulative-Spectrum



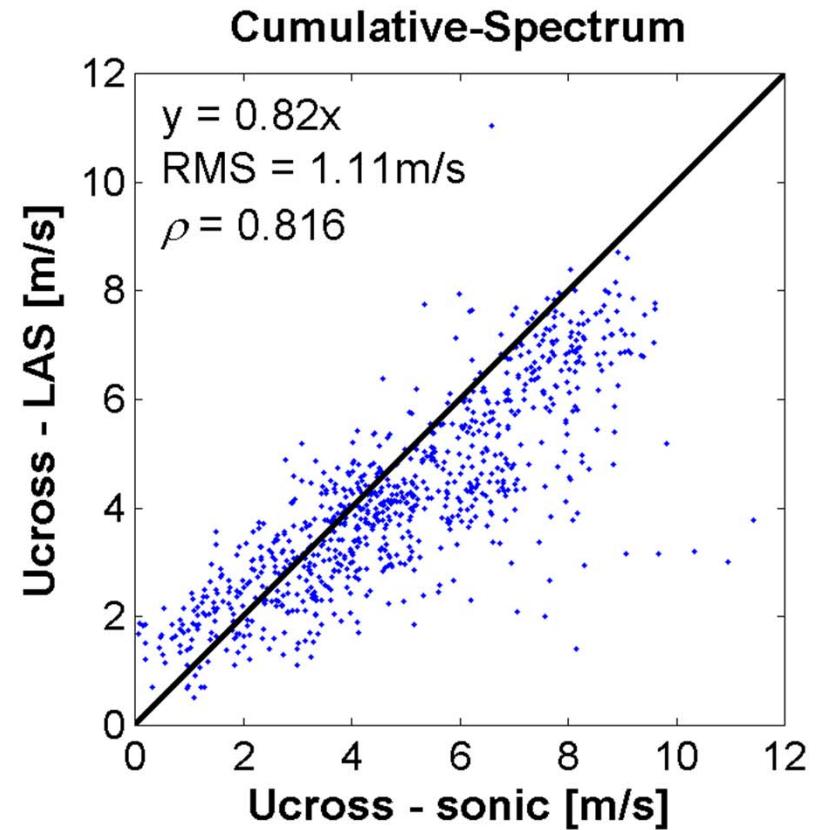
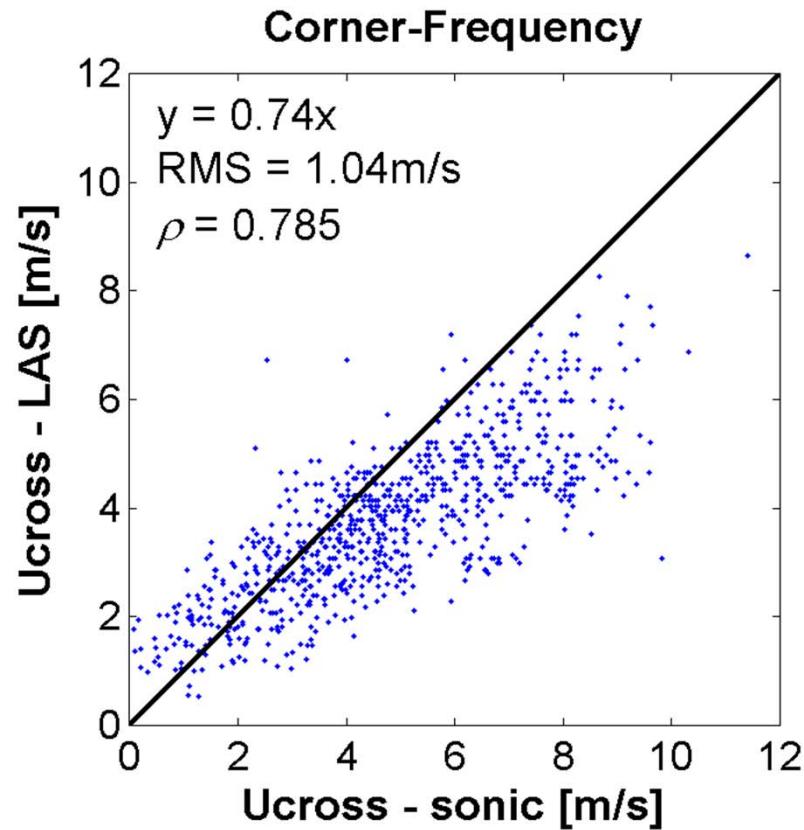
Cross-wind – Results



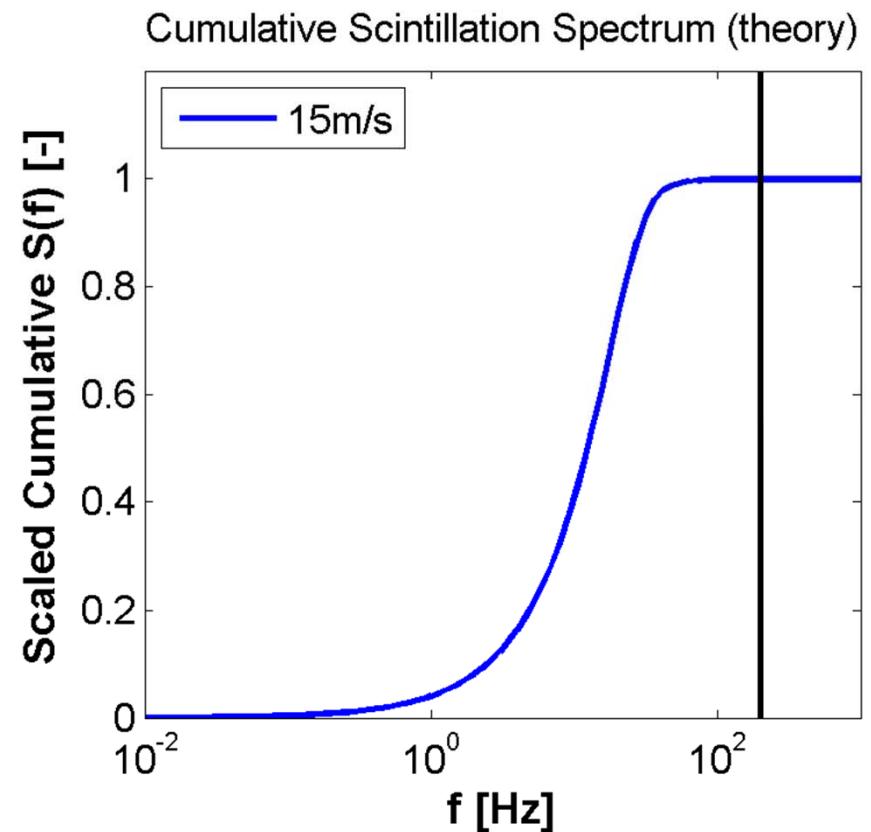
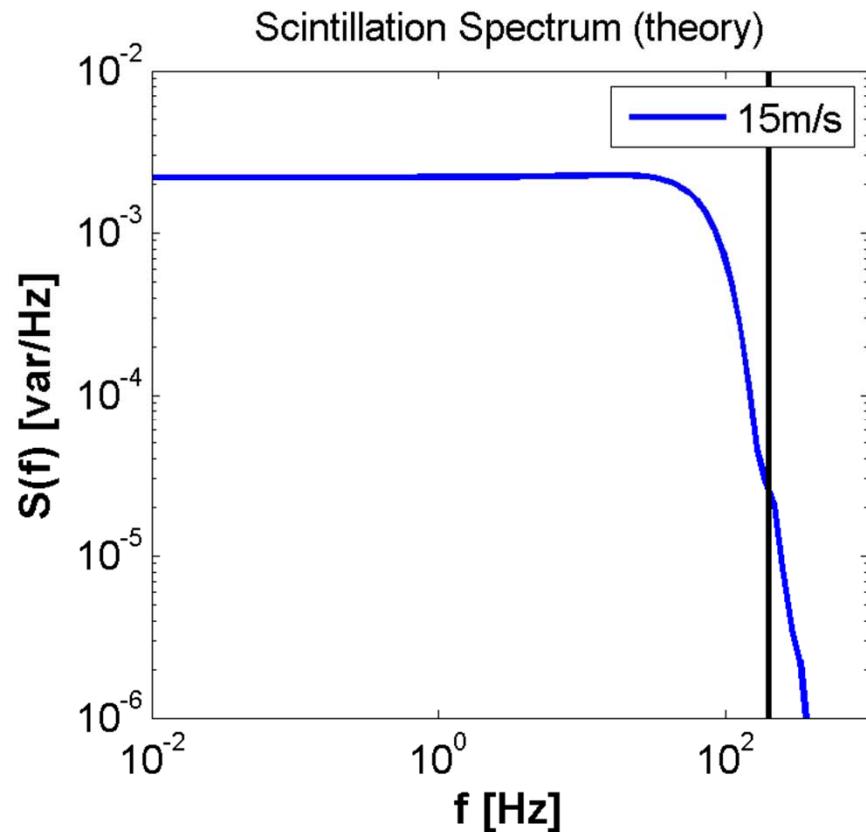
Cross-wind – Results



Cross-wind – Results



Cross-wind – Results



Nyquist frequency @ 250Hz

Conclusions

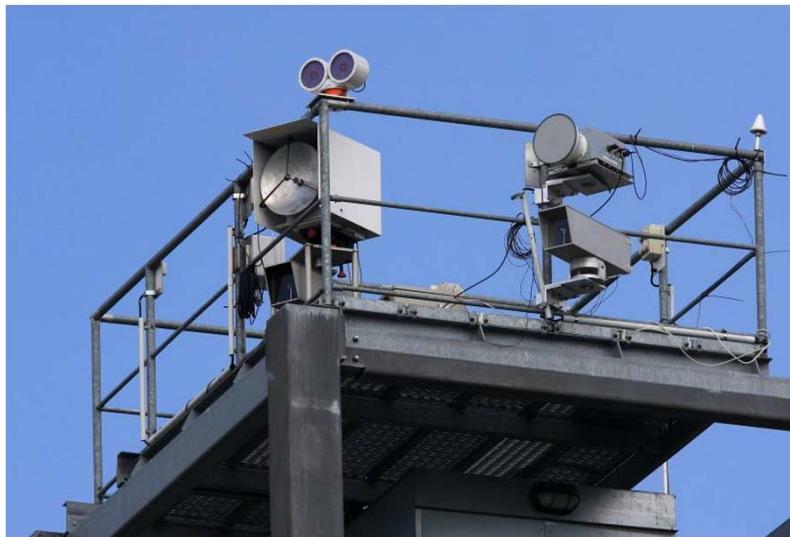
OMS systems:

- still to be analyzed ...

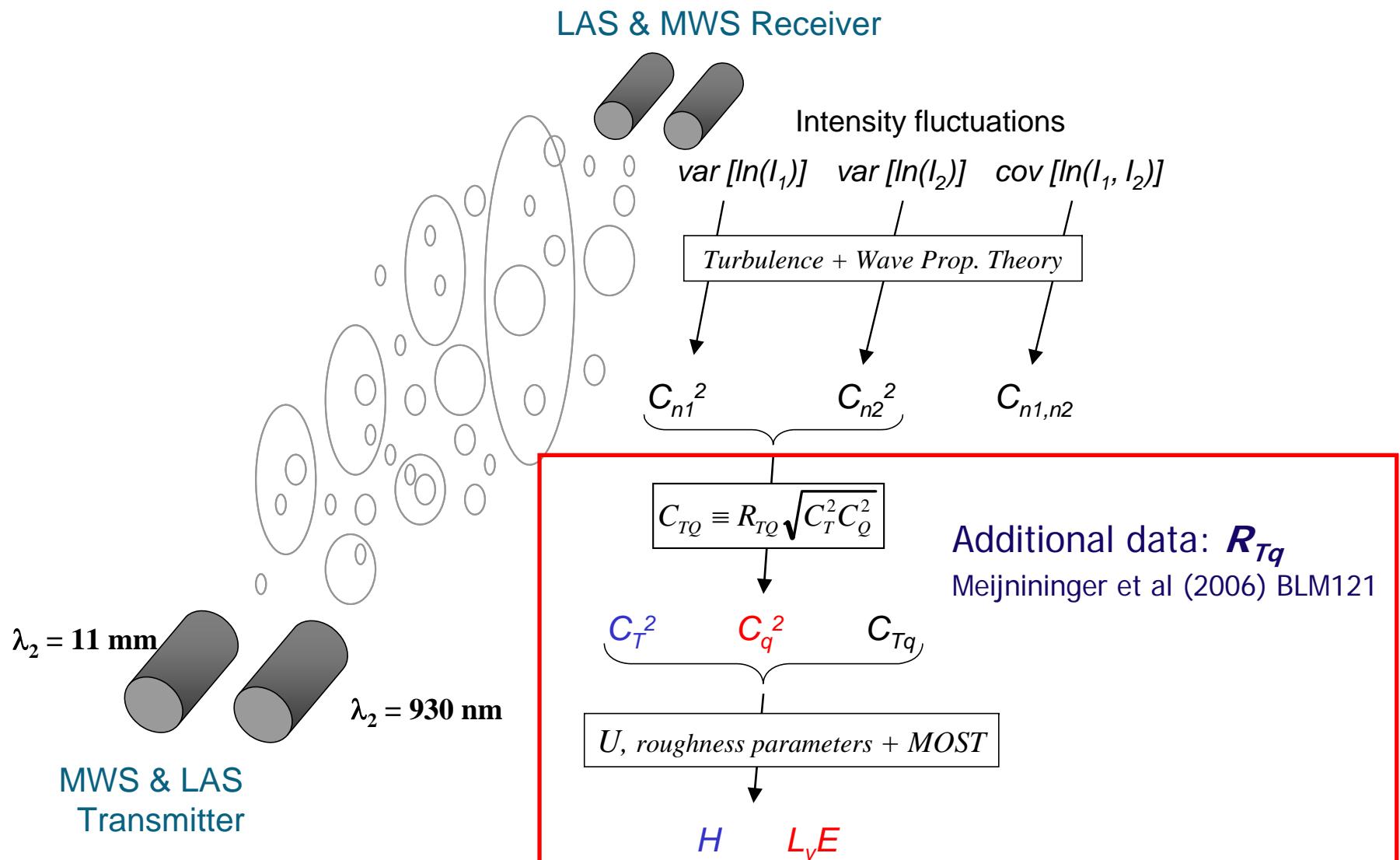
Single Aperture Cross-Wind:

- New approaches are “promising”
- Methods work reasonably well under low wind speed conditions
- Higher sampling rate to improve performance under high wind-speed conditions
- Improve on non-stationary conditions by using local spectra from wavelets

Thank you



OMS system – Third Equation – R_{TQ}



OMS system – Structure parameter of refractive index

Refractive index of air: $n'(\lambda) = A_T(\lambda) \frac{T}{T} + A_Q(\lambda) \frac{Q}{Q} + A_P(\lambda) \frac{P}{P}$

$$A_T(\lambda) \equiv \bar{T} \frac{\partial n(\lambda)}{\partial \bar{T}}$$

$$A_Q(\lambda) \equiv \bar{Q} \frac{\partial n(\lambda)}{\partial \bar{Q}}$$

$$C_{n_\lambda}^2 = \frac{A_{T_\lambda}^2}{\bar{T}^2} C_T^2 + \frac{2A_{T_\lambda} A_{Q_\lambda}}{\bar{T} \bar{Q}} C_{TQ} + \frac{A_{Q_\lambda}^2}{\bar{Q}^2} C_Q^2$$

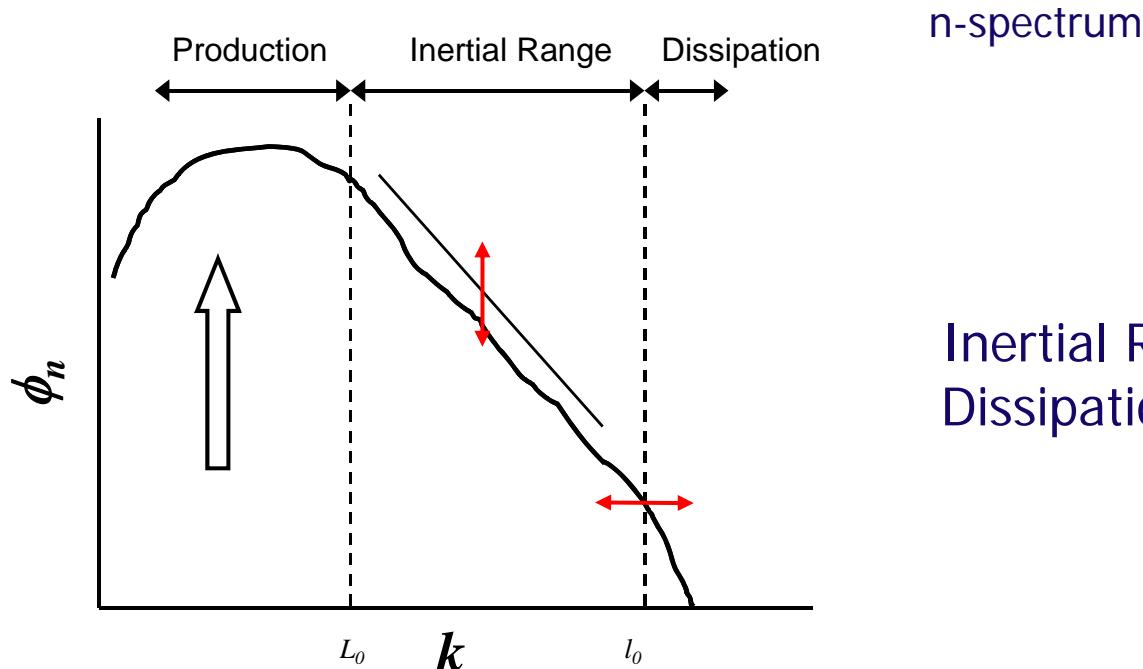
	C_n^2	$\frac{A_T^2}{\bar{T}^2} C_T^2$	$2 \frac{A_T A_Q}{\bar{T} \bar{Q}} C_{TQ}$	$\frac{A_Q^2}{\bar{Q}^2} C_Q^2$
LAS:	$\lambda = 930 \text{ nm}$	1.22×10^{-14}	1.09×10^{-14}	1.19×10^{-15}
MWS:	$\lambda = 11 \text{ mm}$	8.7×10^{-13}	1.7×10^{-14}	-2.49×10^{-13}

H $L_v E$



I Wave Propagation Theory + Turbulence

Scintillometer Equation: $\sigma_{\ln(I)} = 16\pi^2 K^2 \int_0^\infty k \underbrace{\phi_n}_{\text{n-spectrum}} \sin^2 \left[\frac{k^2 x(L-x)}{2KL} \right] \left[\frac{4J_1^2(kDx/2L)}{(kDx/2L)^2} \right] dk dx$

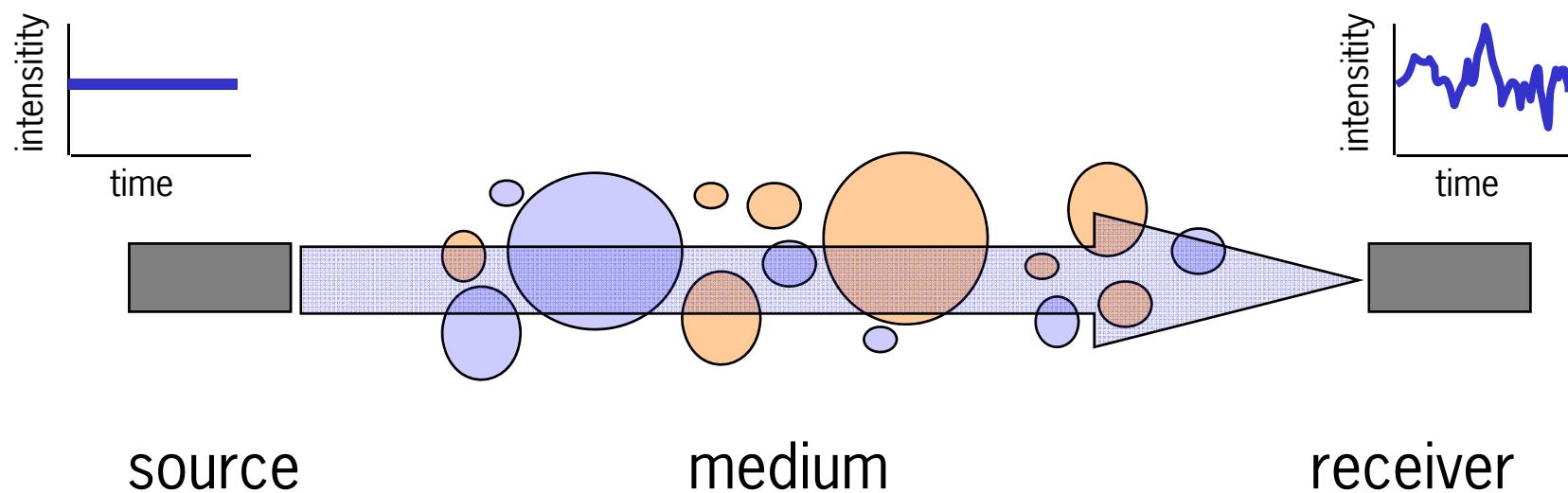
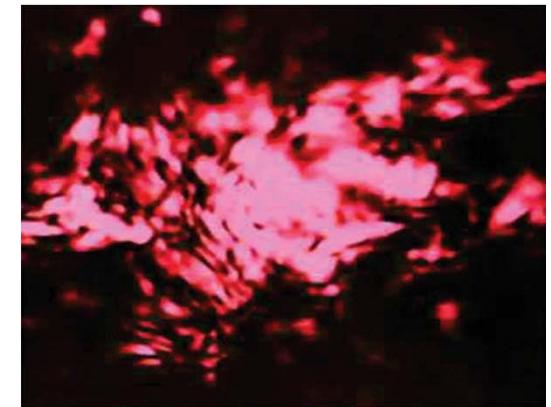


Inertial Range – Kolmogorov
Dissipation Range - Hill

- Scintillometer sensitive to one eddy scale: largest of D or $F = (\lambda L)^{1/2}$
- If in Inertial range $\rightarrow \phi_n = 0.033 \underline{C_n^2} k^{-11/3}$
- If in Dissipation range $\rightarrow \phi_n = 0.033 \underline{C_n^2} \underline{k^{-11/3}} f_A(kl_0)$



Scintillation – Diffraction Process



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Scintillometry

Scintillometer footprint

