# WAGENINGEN UNIVERSITY

# METEOROLOGY AND AIR QUALITY GROUP

MSC THESIS

# Quantifying $NO_x$ emissions from Paris with high-resolution satellite measurements

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#### Abstract

The new high-resolution TROPOMI satellite observations make it now possible to measure nitrogen dioxide (NO<sub>2</sub>) at city level from space. Especially in cities nitrogen dioxides (NO<sub>x</sub>) pose a threat to air quality, since large numbers of people and strong emission sources are concentrated together. This study evaluated and developed a simple column model approach to quantify these NO<sub>x</sub> emissions for Paris from day-to-day with the first TROPOMI observations. The influence of downwind emissions, diurnal variation in emissions, lifetime variability, and background decay on the column model's simulated downwind decay were examined. Because these influences on the downwind plume showed to change the estimated NO<sub>x</sub> lifetime by up to a factor of 3, the NO<sub>x</sub> emission was estimated just from the increase in NO<sub>2</sub> in the wind direction over Paris as measured by TROPOMI. The resulting average estimated NO<sub>x</sub> emission from Paris of eleven days (in November 2017, February and April 2018) was 55.9 mol s<sup>-1</sup> during weekdays and 30.2 mol s<sup>-1</sup> during weekends, which is respectively 24% and 51% smaller than the EDGAR emission inventory from 2012. This research demonstrated how TROPOMI's high-resolution observations can be used to estimate NO<sub>x</sub> for single days. Such an approach can be a fast and understandable tool for evaluating the effectiveness of air quality and climate mitigation measures.

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# 1 Introduction

Atmospheric nitrogen oxides ( $NO_x = NO + NO_2$ ) have an adverse effect on public health and also influence the climate. This is especially an important issue in cities, where both large numbers of people and strong  $NO_x$  emissions from traffic, heating and industry are concentrated. In Paris, for instance, around 1.3 million inhabitants were exposed to an annual mean nitrogen dioxide ( $NO_2$ ) concentration that exceeded the EU limit of 40  $\mu$ g m<sup>-3</sup> in 2017 (Airparif, 2018). To comply with the EU air quality norms by 2025, the city of Paris is currently planning 46 new mitigation measures in their latest atmospheric protection plan (DRIEE, 2018). These measures range from stimulating cycling to imposing more stringent emission norms on industry, but also include additional actions during strong pollution episodes.

Such air quality and climate mitigation plans require monitoring of  $NO_x$  emissions to evaluate their effectiveness. This emission monitoring is often done with a bottom-up approach, where the emission is calculated based on activity data and corresponding emission factors. However, there are substantial uncertainties in this method (Crippa et al., 2018; Kuenen et al., 2014), for instance because the exact location of emission is not known. That is why these bottom-up emission inventories are validated with measurements, the so-called top-down approach, often in combination with a model that relates observed atmospheric concentrations to emissions. Such a top-down study based on MAX-DOAS observations by Shaiganfar et al. (2017) found between 1.4 and 2.3 times higher  $NO_x$  emissions from Paris compared to bottom-up emission inventories.

Also top-down studies that use satellite observations of  $NO_2$  have resulted in important insights into  $NO_x$  emission sources. Because the spectrometers on these satellites measure reflected solar radiation from which tropospheric  $NO_2$  columns are retrieved, they grasp all the  $NO_x$  that has been emitted. For example, Akimoto et al. (2006) showed that the coal consumption of China was significantly under reported from 1996 to 2003. In addition, temporally averaged remote sensing observations of  $NO_2$  have been used to quantify emissions of isolated sources such as ship tracks (Vinken et al., 2014) and megacities (Beirle et al., 2011).

On 13 October 2017, a new satellite with the TROPOspheric Monitoring Instrument (TROPOMI) was launched, a spectrometer that allows observation of key atmospheric compounds, including  $NO_2$  (Veefkind et al., 2012). This new instrument provides daily global measurements at a resolution of  $3.5 \times 7$  km<sup>2</sup>, which is more than three times finer than its predecessor OMI. With approximately 130 pixels over Paris, TROPOMI makes it possible to examine tropospheric  $NO_2$  columns at city level. This raises the question how these new high-resolution  $NO_2$  column observations can allow a better quantification of  $NO_x$  emissions.

This study aims to evaluate and develop a simple column model approach to quantify the  $NO_x$  emissions of Paris from single-day TROPOMI observations. The column model, as explained by Jacob (1999), describes the chemical evolution of a pollutant in the wind direction as a function of the emission, chemical decay and wind speed. The emission estimates from the column model are compared to the bottom-up EDGAR emission inventory. In chapter 4, this method is introduced for 22 November 2017, one of the first TROPOMI observations at cloud-free conditions over Paris, showing a distinct  $NO_2$  plume. Different factors that influence the decay downwind of Paris are further evaluated in chapter 5. Since these downwind influences on the column model estimates of the  $NO_x$  lifetime were shown to be large, the last part of this study, chapter 6, estimates  $NO_x$  emissions just from the increase in  $NO_2$  along the wind direction over the source area. This is done for ten selected days in the months February to April 2018 and optimally shows the new possibilities of TROPOMI's higher resolution

# 2 Background

## 2.1 Tropospheric nitrogen dioxides

According to the World Health Organization, exposure to  $NO_2$  is linked to a reduced long function and increased bronchitis symptoms of asthmatic children (WHO, 2016). Also the ozone ( $O_3$ ) that is formed from  $NO_2$  has a negative effect on human health; it can cause breathing problems and lung diseases and is associated with increasing mortality rates. Besides health impacts,  $O_3$  and hydroxyl (OH) radicals formed from tropospheric  $NO_2$  also have an impact on climate: the former acts as a greenhouse gas and the latter enhances the oxidation and shortens the lifetime of methane ( $CH_4$ ), the second most significant anthropogenic greenhouse gas. Finally, the deposition of  $HNO_3$  causes fertilisation of soils and surface waters.

Although almost all  $NO_x$  is emitted as nitrogen oxide (NO), there is a rapid cycling between NO and  $NO_2$  in the atmosphere (Jacob, 1999). NO reacts with  $O_3$  to form  $NO_2$  and the reverse reactions is possible in the presence of oxygen and solar radiation. Fossil fuel combustion accounts for about half of the  $NO_x$  present in the troposphere, while biomass burning, mainly from tropical agriculture and deforestation, comprises another quarter (Jacob, 1999, p. 212).

In the troposphere,  $NO_x$  has a short lifetime, in the order of one day, mainly because of the oxidation of  $NO_2$  to  $HNO_3$  after which  $HNO_3$  is scavenged by precipitation. During daytime  $NO_2$  is oxidised by OH, whereas at night  $HNO_3$  is formed through  $NO_3$  and  $N_2O_5$ . A smaller sink is the oxidation of  $NO_2$  to peroxyacetylnitrate (PAN), which is enabled by the presence of photochemically oxidised hydrocarbons. The lifetime of PAN is strongly dependent on temperature, so that PAN that is present in the middle and upper troposphere can be transported over long distances before it decomposes back to  $NO_x$ .

Tropospheric  $NO_x$  also has an important role in the cycle of OH. The photo-chemical processing of  $NO_2$  results in the production of NO and  $O_3$ , which can photolyse and produce OH molecules that enable the oxidation of CO and  $CH_4$ .

# 2.2 Emission monitoring

Bottom-up emission inventories, based on activity data and corresponding emission factors, have large uncertainties in the reported  $NO_x$  emissions. In the TNO-MACC emission inventory, the uncertainty ranges from 20% up to 300% for different source categories (Kuenen et al., 2014). In the Atmospheric Database for Global Atmospheric Research (EDGAR) v4.3.2 inventory of 2012, which is used in this research, the uncertainty in  $NO_x$  emission from different regions varies from 17.2% to 69.4% (Crippa et al., 2018). And specifically for megacities, Butler et al. (2008) found that three different emission inventories often differ by a factor of two for the same city, which could possibly be explained by the fact that the spatial allocation of emissions is based on population density rather than the actual point of emission.

Top-down studies on  $NO_x$  emissions based on satellite measurements have been an important tool to evaluate bottom-up accounting. This satellite monitoring of  $NO_x$  emissions can be distinguished in three different approaches:

- 1. Formal inversions using chemistry transport models (e.g. Martin et al., 2003)
- 2. Comparisons between the change in reported  ${\rm NO}_x$  emissions and observed  ${\rm NO}_2$  columns (e.g. Jiang et al., 2018)

#### 3. A column model for single source areas (e.g. Beirle et al., 2011)

Although the total global  $NO_x$  emission estimated with the top-down approach based on satellite observations in combination with atmospheric models generally align closely with bottom-up reported global annual emissions (e.g. Martin et al., 2003), these top-down studies often indicate large discrepancies in the magnitude, temporal variation and regional distribution of emission sources compared to bottom-up inventories. In the United States,  $NO_2$  satellite measurements of OMI show a slower reduction than the predicted  $NO_x$  emission trend of the Environmental Protection Agency's inventory from 2011 to 2015 (Jiang et al., 2018). Martin et al. compared emissions derived from GOME and an inverse model with the GEIA and EDGAR 3.0 emission inventory and found significant regional differences (2013). Richter et al. found a stronger positive trend in the tropospheric  $NO_2$  concentration over China in satellite observations than bottom-up inventories (2005). Another study on  $NO_x$  emissions in China that used the different passing time of GOME-2A and OMI found clear local differences (Lin et al., 2010). Finally, satellite observations also enable studies on the weekly or seasonal cycle of  $NO_2$  (e.g. Beirle et al., 2003). These local and temporal differences can have an important influence on the accuracy of modelling studies that use bottom-up emission inventories, especially because of the short lifetime of  $NO_x$ .

Besides top-down studies at global or continental scales, temporally averaged satellite observations have been used to directly estimate emissions from a single emission source. Vinken et al. found that shipping emissions are both overestimated (up to 60%) and underestimated (up to 131%) in comparison to the EMEP inventory (2014). Another example is a study by Wang et al., which used OMI observations and a global chemical transport model to investigated the contribution of new coal power plants in China to  $NO_x$  emissions (2012).

Rather than using a full chemical transport model,  $NO_x$  emissions from a single source area have also been successfully estimated with a column model similar to the method described in Jacob (1999). De Foy et al. evaluated different implementation of this approach to estimate surface emissions and lifetimes from satellite measurements (2014). Beirle et al. has applied this method successfully to estimated  $NO_x$  emissions and lifetimes from ship tracks (2004), as well as from megacities (2011). This showed generally a good agreement with the EDGAR inventory, but an underestimation of 300% for the city of Riyadh.

However, these column model studies require averaged data for multiple years to enhance the spatial resolution. That is why the advancement in resolution of TROPOMI (described in chapter 3) is particularly promising for estimating  $NO_x$  emissions with this third satellite emission monitoring approach. This could enable a quantification of daily emissions rather than annual averages based on a single TROPOMI orbit.

# 3 Data description

#### 3.1 TROPOMI observations

Currently, four satellites are orbiting the planet to monitor tropospheric  $NO_2$ : GOME-2A, GOME-2B, OMI and its successor TROPOMI (Boersma et al., 2018). The GOME-2 instrument passes the equator at 10:30 local time, while the OMI and TROPOMI both have an early afternoon overpass of respectively 13:40 and 13.30 local time (Boersma et al., 2018; Veefkind et al., 2012). The precise TROPOMI overpass times at Paris for the selected days are listed in Table 3.1.

The tropospheric  $NO_2$  column is retrieved from the spectrometer measurements of direct and backscattered solar radiation with an algorithm, such as DOMINO v2.0 (Boersma et al., 2011) and SP2 (Buscela et al., 2013) for OMI. An update of the DOMINO v2.0, the new QA4ECV product, is now available for the OMI and GOME-2A observations, which started respectively in 2004 and 2007, as well as for past measurements by GOME-2A, from 1996 to 2003, and SCIAMACHY, from 2000 to 2012 (Boersma et al., 2018). For TROPOMI, a retrieval algorithm similar to QA4ECV is used (van Geffen et al., 2016). The general functioning of this retrieval procedure can be described in three steps (ibid.):

- 1. Derive the NO<sub>2</sub> slant columns from the measured radiance and irradiance spectra with the Differential Optical Absorption Spectroscopy (DOAS) method.
- 2. Separate the tropospheric and stratospheric component of the slant columns.
- 3. Convert the tropospheric and stratospheric slant columns to a vertical column based on the tropospheric air mass factor (AMF).

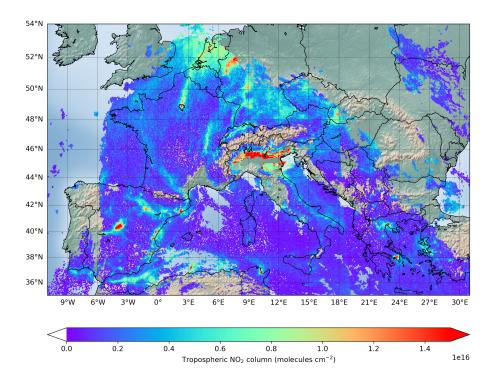


Figure 3.1: Tropospheric NO<sub>2</sub> columns observed by TROPOMI on 22 November 2017 over Europe with AMF, albedo and cloud fraction filters.

The NO<sub>2</sub> tropospheric column is observed with a spatial resolution of  $3.5 \times 7 \text{ km}^2$ , much higher than OMI's resolution of  $13 \times 24 \text{ km}^2$  (Boersma et al., 2011). In comparison to its predecessor OMI, the new TROPOMI spectrometer (described by Veefkind et al., 2012) has an extended wavelength range with bands in the NIR and SWIR in addition to the UV and VIS range. This new NIR band provides information about cloud characterisation and enables a better cloud correction, resulting in a smaller error for partly cloudy conditions. In addition, it is expected that the smaller pixel size of TROPOMI observations leads to 70% more cloud-free retrievals in comparison to OMI, since even a small clouds affects the whole pixel (Krijger et al., 2007). Furthermore, TROPOMI's signal-to-noise ratio is a factor 2-3 higher than OMI.

In chapter 4 of this study, preliminary TROPOMI data of 22 November 2017 is used, with cloud free conditions over central Europe, showing the tropospheric NO<sub>2</sub> columns and a distinct north-northeasterly plume originating from Paris (Figure 3.1). For the last section, chapter 6, TROPOMI observations of two periods were analysed: 22 to 26 February 2018 and 17 to 24 April 2018. This preliminary data is from the commissioning phase and hence a good test of the usefulness of the early-mission data quality before the public data release on 11 July 2018. TROPOMI data is now available via the Copernicus Open Access Hub<sup>1</sup>.

## 3.2 EDGAR emission inventory

To compare the emissions that are derived from the TROPOMI observations, the bottom up emission inventory EDGAR v4.3.2<sup>2</sup> of 2012 is used as reference (Figure 3.2). This version of the EDGAR emission inventory is described by Crippa et al. (2018). The EDGAR emissions in EU28 countries, which includes Paris, are reported to have an uncertainty of 50.7% (ibid.). These reported emissions are based on activity data and spatial distribution proxies from Janssens-Maenhout et al. (2017), similar to the EDGAR greenhouse gas emission inventory, and emission factors from the EEA Guidebook (2013), HTAP\_v2 (Janssens-Maenhout et al., 2015) and PEGASOS (Crippa et al., 2016).

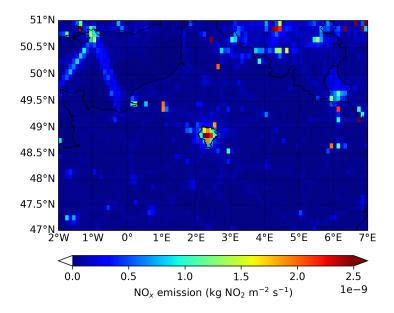


Figure 3.2:  $NO_x$  emissions, expressed in  $NO_2$ , from the EDGAR inventory from 2012 for Paris and its surroundings.

<sup>&</sup>lt;sup>1</sup>https://scihub.copernicus.eu/

<sup>&</sup>lt;sup>2</sup>https://data.europa.eu/doi/10.2904/JRC\_DATASET\_EDGAR

## 3.3 Radiosondes in Trappes

Every day at noon a sounding is performed by Meteo France at Trappes, approximately 30 km West-Southwest of the centre of Paris, which gives the vertical profile of, among other variables, the wind speed and direction (Table 3.1). The boundary layer height is determined from the inversion in the potential temperature profile (Table 3.1). These radiosonde observations are available via the University of Wyoming<sup>3</sup> and the surface observations via the National Oceanic and Atmospheric Administration<sup>4</sup>. It should be noted that the site is located 168m above sea level, whereas the centre of Paris is approximately 130 meters lower.

#### 3.4 Paris surface observations

Surface NO and NO<sub>2</sub> observations in the Paris area are available from Airparif<sup>5</sup>. Air is analysed using catalytic reduction and chemiluminescence according to the European EN 14211 standard (Jordan Bureau, personal communication, 24 August 2018). The hourly averages for 40 observations sites that measure NO<sub>2</sub> concentrations in  $\mu$ g m<sup>-3</sup> were available, of which 13 traffic stations were excluded because of strong local contributions from traffic and 4 outside the domain of the source area. The location of the stations is shown in the maps in Figure 6.2 and 6.3. In addition, NO and NO<sub>2</sub> concentrations are measured at the top of the Eiffel Tower at a height of 315 m, which gives important information on the vertical distribution and is more representative for the whole city rather than surface stations that measure concentrations influenced by emissions at street level.

## 3.5 Day selection criteria

Not all days have TROPOMI observations that are suited to estimate emissions with the column model approach. The main limitation is that only days which are largely cloud fee can be used, since clouds screen the lowest part of the atmosphere where  $NO_x$  emission takes place. Overall, three conditions were considered for the day selection in this study:

- 1. Cloud-free conditions are required to interpret the increase and decay of NO<sub>2</sub> along with the wind.
- 2. No problems in the tropospheric AMF, which is tested by dividing the difference between the slant column density and the stratospheric slant column density by the geometric AMF.
- 3. A sufficiently strong wind speed so that an increase in  $NO_2$  over the source area in the wind direction can be observed, from which the  $NO_x$  emission can be determined with the column model.

Chapter 4 uses TROPOMI observations from 22 November 2017, as this was one of the first available days showing a clear plume from Paris (Figure 3.3a). From the period February to April 2018, two sequences of days were cloud free: 22 to 26 February and 17 to 24 April. The orbit of 24 February 2018 was excluded because of problems in the tropospheric AMF. Two more days in April 2018 were excluded: the 20th because of strong stagnant and partly cloudy conditions shown by little outflow of  $NO_2$  out of the source area (Figure 3.3b) and the 23rd because of a high cloud fraction over Paris. This results in a total of eleven days that are used.

<sup>&</sup>lt;sup>3</sup>http://weather.uwyo.edu/upperair/bufrraob.shtml

<sup>&</sup>lt;sup>4</sup>https://www.ncdc.noaa.gov/isd

 $<sup>^5 {\</sup>rm https://www.Airparif.asso.fr/stations}$ 

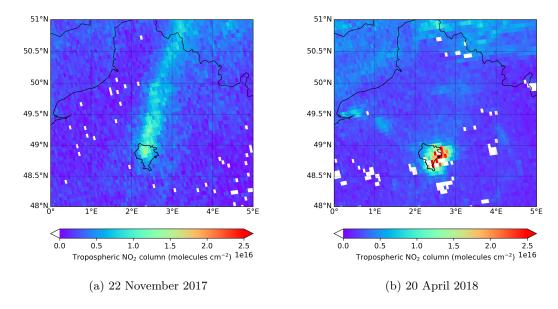


Figure 3.3: Tropospheric NO<sub>2</sub> columns observed by TROPOMI over Paris (denoted by the black outline) and its surrounding region with AMF, albedo and cloud fraction filters.

#### 3.6 First validation of TROPOMI over Paris

In principle, one could calculate an NO<sub>2</sub> column by integration the concentration measured at the Eiffel Tower up to the boundary layer height, assuming that the boundary layer is well-mixed and that the NO<sub>2</sub> concentration is constant with height. However, the results of this approach deviate substantially with the tropospheric NO<sub>2</sub> columns measured with TROPOMI (Table 3.1). Only for three days out of ten days the Eiffel Tower column and TROPOMI column differ by less then 25%. Except for 22 November 2017, all the Eiffel Tower columns are substantially higher. The Eiffel Tower column is even a factor 5 higher than the TROPOMI tropospheric column on 24 February 2018, a day with a relatively high boundary layer height of 2000 m.

The higher Eiffel Tower column suggests that NO<sub>2</sub> is not fully mixed within the boundary layer and that the concentration is higher at the surface. This is also shown by NO<sub>2</sub> measurements with a ground-based UV-visible light spectrometer in the centre of Paris (Dieudonné et al., 2013). Such a vertical NO<sub>2</sub> gradient is also expected, since NO<sub>2</sub> is continuously emitted at the surface. Thus assuming that the Eiffel Tower measurements are representative for the full boundary layer leads to an overestimation of the column.

Table 3.1: Overview of all selected days, with TROPOMI overpass time at Paris, boundary layer height (h), as well as wind speed (U) and direction in the middle of the boundary layer from sounding observations at Trappes. Surface wind speed and direction at Trappes are also given. Wind data marked with an asterisk is the boundary layer mean from the ECMWF-Interim reanalysis (unpublished paper, Lorente et al., 2018). TROPOMI tropospheric NO<sub>2</sub> column are given in comparison to the calculated NO<sub>2</sub> column based on Eiffel Tower measurements and h.

				Boundary	y layer	Surface		Column	
Date	Day	Overpass	h	U	Direction	U	Direction	TROPOMI	Eiffel Tower
		(UTC)	(m)	$(\mathrm{m}\ \mathrm{s}^{-1})$	(degrees)	$(\mathrm{m}\ \mathrm{s}^{-1})$	(degrees)	$(\cdot 10^{16} \ \mathrm{NO_2} \ \mathrm{r}$	$\rm molecules~cm^{-2})$
22/11/17	Wednesday	12:52	500	8.8	185	4.6	190	1.5	1.2
22/02/18	Thursday	12:26	874*	8.6*	50*	5.7	50	1.2	3.1
23/02/18	Friday	12:07	907*	9.0*	60*	7.2	60	0.9	2.4
24/02/18	Saturday	13:30	2000	5.7	26	7.2	50	0.9	4.7
25/02/18	Sunday	13:11	1000	9.3	54	6.7	40	0.7	1.3
26/02/18	Monday	12:52	1400	9.8	50	8.2	60	0.9	2.2
17/04/18	Tuesday	12:12	1250	5.7	186	3.6	200	0.8	1.0
18/04/18	Wednesday	11:53	1350	8.2	99	5.1	120	0.9	3.2
19/04/18	Thursday	13:16	750	4.1	194	4.1	220	0.9	1.1
20/04/18	Friday	12:57	1700	4.6	289	3.1	10	1.1	4.2
21/04/18	Saturday	12:38	1500	4.1	83	3.1	150	1.1	2.7
22/04/18	Sunday	12:19	2188*	9.7*	230*	4.6	230	0.4	1.1
24/04/18	Tuesday	13:23	1200	4.6	209	4.1	220	0.2	-

# 4 Applying the column model

The TROPOMI observations from 22 November 2017, illustrated in Figure 3.3a, show accumulation of  $NO_2$  over Paris along with the wind followed by a decaying plume downwind of the source area. From these measurements, a one-dimensional line density is calculated in this chapter, which gives the evaluation of the number of  $NO_2$  molecules at a certain distance from the source area along the wind direction. Subsequently, the  $NO_x$  emission rate can be estimated with a column model similar to Jacob (1999) from this line density and minimal additional model input. This study focuses on the city of Paris, as it has a substantial source area, only limited emission sources in its surrounding, and one of the first TROPOMI observations shows a distinct plume from Paris on 22 November 2017.

## 4.1 Methodology

#### 4.1.1 Line density

The tropospheric  $NO_2$  columns from the TROPOMI observations are selected as a function of the distance from the source area of Paris in the wind direction (x) and integrated perpendicular to the wind direction (y), similar to the approach used by Beirle et al. (2011). This line density represents the change in the  $NO_2$  column along with the wind, showing a steep increase in  $NO_2$  as the column moves over Paris and a slower decay as the column moves away from the source area. The wind direction in the middle of the boundary layer is selected from the radiosonde measurements at Trappes.

The line density starts 30 km from the centre of Paris and continues in the direction from the wind, see Figure 4.1. The source area runs from 0 to 60 km on the line density, which represents approximately the greater Paris area (Métropole du Grand Paris). The line density continues up to 140 km downwind of the source area. Although it appears from the TROPOMI observation that the plume is still visible more than 140 km downwind on 22 November 2017, this boundary is chosen to avoid mixing up the plume from Paris with local enhancements in NO<sub>2</sub> from fresh emissions, since the EDGAR emission inventory shows some relatively large emission sources further downwind close to the French-Belgium border (Figure 3.2).

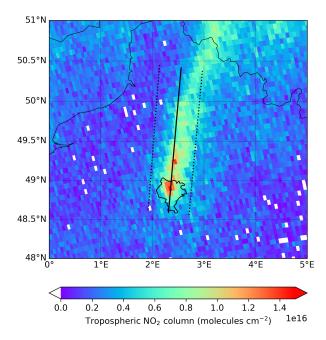


Figure 4.1: Tropospheric  $NO_2$  columns observed by TROPOMI on 22 November 2017. The greater Paris region, integration line and across track integration boundaries are delineated by the black lines. For the white pixels, the measurements have been excluded based on the AMF, albedo and cloud fraction criterion.

For every interval of 3.5 km (similar to TROPOMI's resolution) on the line that passes through the centre of Paris (48°51'53.0"N 2°20'56.5"E) in the wind direction, the tropospheric NO<sub>2</sub> column  $(N_{NO_2})$  pixels at a location (x, y) that are perpendicular to and within a distance of 30 km from the centre line (y = 0) are selected. Both the centre line in the x direction and the perpendicular lines in the y direction are illustrated in Figure 4.2. The total source area is hence 60 by 60 km<sup>2</sup>, tilted in the direction of the wind and with the centre of Paris in the middle. The selected columns are multiplied by the intersection length  $(\delta y)$  of the borders of the pixel with the line perpendicular to the wind direction to obtain the line density for each selected pixel (Figure 4.2, where the × markers denote the intersections). This intersection length can vary from very small up to the length of the diagonal of the pixel (7.8 km). The selected pixel line densities  $(N_{NO_2}(x, y) \cdot \delta y)$  are subsequently summed, resulting in the total line density in NO<sub>2</sub> molecules cm<sup>-1</sup> at x:

$$L_{NO_2}(x) = \sum_{-30 < y < 30} N_{NO_2}(x, y) \cdot \delta y \tag{4.1}$$

Measurements with a tropospheric air mass factor (AMF) lower than 0.2, surface albedo higher than 0.3 and cloud radiance fraction in the  $NO_2$  window higher than 0.5 are excluded. These excluded pixels are replaced with the average of the line density at this distance, for instance for the pixel in the lower left corner in Figure 4.2. If half of the across length does not meet these selection criteria, then no line density is given for that interval since the line density would not be representative for entire domain length.

4.1. Methodology

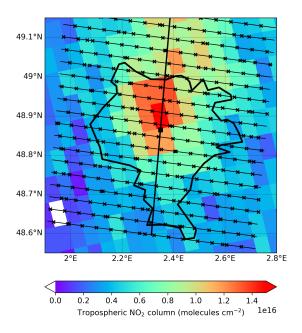


Figure 4.2: Line density derivation and tropospheric NO<sub>2</sub> columns over the greater Paris region on 22 November 2017. The centre line starts 30 km from the centre of Paris (denoted by the square) and proceeds in the direction of the wind (x). For each of the lines at distance x perpendicular to the wind direction (y) a line density is calculated, which is the sum of all the pixels on this line multiplied by the intersection length of the respective pixel with this line  $(\delta y)$ . The intersection length is illustrated as the distance between the  $\times$  markers in this figure.

#### 4.1.2 Column model to simulate line densities

The column model, described in Jacob (1999), simulates the chemical evolution of a pollutant in a well-mixed column of air moving in the direction of the wind with speed U, passing a source area with length  $x_s$  with a constant emission rate E of  $NO_x$  in molecules  $cm^{-1}$  s<sup>-1</sup>. Here the column has the width of the plume of 60 km as defined in the line density derivation. Since most of the tropospheric  $NO_2$  is present in the boundary layer, the height of the modelled column represents this layer. The wind speed for the middle of the boundary layer is derived from the sounding observations at Trappes. The background line density is represented by the constant B.

A first-order chemical loss is assumed with a rate constant k of  $NO_x$ , which is inversely proportional to the effective lifetime of  $NO_x$  ( $\tau_{NO_x}$ ) in the plume. During the middle of the day, this loss is primarily determined by the reaction of  $NO_2$  with OH to HNO<sub>3</sub> (Seinfeld and Pandis, 2006):

$$NO_2 + OH + M \xrightarrow{k_2} HNO_3 + M$$
 (R4.2)

The  $NO_x$  lifetime in hours can be calculated as a function of the rate constant of reaction R4.2, the OH concentration and the  $NO_x/NO_2$  ratio:

$$\tau_{NO_x} = \frac{1}{k} = \frac{1}{k_2[OH]} \cdot \frac{[NO_x]}{[NO_2]}$$
(4.3)

The line density of  $NO_x$  at distance x along the wind direction is then given by the following two formulas:

$$L_{NO_x}(x) = \frac{E}{k} \left( 1 - e^{-\frac{kx}{U}} \right) + B \text{ for } 0 \le x \le x_s$$

$$\tag{4.4}$$

$$L_{NO_x}(x) = L_{NO_x}(x_s)e^{-\frac{k(x-x_s)}{U}} + B \text{ for } x > x_s$$
 (4.5)

Equation 4.4 describes the accumulation of  $NO_x$  as the column passes the source area with length  $x_s$  in the main wind direction. The first term on the right-hand side of this equation denotes the  $NO_x$  emission strength over the full width of the column (60 km) proportional to the rate constant. The second term on the right-hand side of equation 4.4 describes the exponential chemical decay (reaction 4.2) as a function of x, which becomes the only factor once the column moves away from the emission source as described in equation 4.5. Thus the steepness of the decay in the line density together with the average wind speed is a measure for the chemical conversion that determines the  $NO_x$  lifetime. This model assumes a constant background, a uniform wind speed with height and in time, zero downwind emissions and no temporal variability in the emission and loss rate of  $NO_x$ .

#### 4.1.3 Estimating emissions with the column model

The column model can not only be used to simulate a line density, but also to estimate emissions as well as the lifetime and background line density. Yet it is important to first account for the difference in units when converting from line density to emission: TROPOMI measures the number of  $NO_2$  molecules, while the emission describes the amount of  $NO_x$ . Here an average correction factor is applied for the  $NO_x/NO_2$  ratio of 1.32, a typical value for urban conditions around noon, similar to Beirle et al. (2011) based on Seinfeld and Pandis (2006):

$$\frac{[NO_x]}{[NO_2]} = \frac{L_{NO_x}}{L_{NO_2}} = \frac{\tau_{NO_x}}{\tau_{NO_2}} \approx 1.32 \tag{4.6}$$

This average value is compared to the ratio of the  $NO_x$  and  $NO_2$  mixing ratios as observed at the 300 m level of the Eiffel Tower. If the boundary layer is well-mixed, then these measurements are representative for the whole layer, whereas surface observations would be perturbed by nearby emissions.

To obtain top-down estimates of E, k and B, equation 4.4 and 4.5 are simultaneously fitted to the TROPOMI line density based on a non-linear least squares regression. This approach assumes that E is constant in time and over the city and that  $\tau_{NO_x}$  can be estimated with a single time-independent value. For a wide range of values of E, k and B, numerous line densities are simulated  $(L_m)$  from which the residuals with n TROPOMI line densities  $(L_o)$  over the standard error of the TROPOMI line density  $(\sigma)$  are calculated with the Levenberg-Marquardt algorithm. The optimal parameters and corresponding covariances from which the standard deviation is derived are selected from the simulated line density that results in the lowest chi squared  $(\chi^2)$ :

$$\chi^{2} = \sum_{i=1}^{n} \left( \frac{L_{o}(x_{i}) - L_{m}(x_{i}, E, k, B)}{\sigma(x_{i})} \right)^{2}$$
(4.7)

The standard error of each line density at  $x_i$  is given by equation 4.8, which accounts for the different relative contributions of each pixel  $N(x_i, y)$  with intersection length  $\delta y$  to the total line density  $L(x_i)$  with length  $\Delta y$  (60 km):

$$\sigma(x_i) = \sqrt{\sum_{-30 < y < 30} \left(\frac{\delta y(x_i, y)}{\Delta y} \cdot \left(N(x_i, y) - \frac{L(x_i)}{\Delta y}\right)\right)^2}$$
(4.8)

4.2. Results

In order to compare the resulting top-down emission estimate from the TROPOMI line density, the EDGAR  $NO_x$  emissions in kg s<sup>-1</sup> are expressed as  $NO_2$  and therefore converted to mol s<sup>-1</sup> with the molecular weight of  $NO_2$ .

#### 4.2 Results

The line density from the TROPOMI observations on 22 November 2017 shows a steep increase in NO<sub>2</sub> over Paris (x < 60 km), while the amount of NO<sub>2</sub> decreases slower downwind of the source area due to chemical decay (Figure 4.2). Subsequently, the emission, lifetime and background concentration is estimated by fitting the column model (Function 4.4 and 4.5) to this line density. The simulated line density leading to the best match with the observed line density has a NO<sub>x</sub> emission of 74.6  $\pm$  4.58 mol s<sup>-1</sup>, lifetime of 3.7  $\pm$  0.28 hours and background of 1.6  $\pm$  0.09  $\cdot$ 10<sup>22</sup> molecules cm<sup>-1</sup> (Table 4.1). This emission estimate from the TROPOMI line density on 22 November 2017 is 38% higher than the EDGAR NO<sub>x</sub> emission inventory from 2012, which gives a value of 54.1 mol s<sup>-1</sup> for the same source domain. The non-linear least squares fit has an R<sup>2</sup> value of 0.55 compared to the TROPOMI line density.

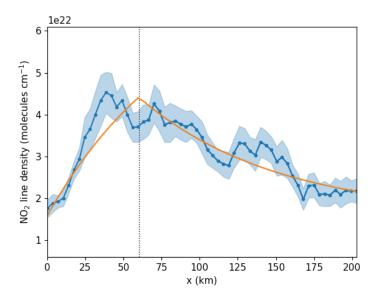


Figure 4.3:  $NO_2$  line density from the TROPOMI observations on 22 November 2017 (blue) and the line density from the inverse column model (orange). The dotted line indicates the length of the source area and the shaded area shows the standard error of the line densities.

Table 4.1: Estimates of the  $NO_x$  emission, lifetime and  $NO_2$  background line density and their standard deviations from fitting the column model to the TROPOMI line density of 22 November 2017.

Parameter	Estimate	Standard deviation	Unit
Emission	74.6	4.68	$\mathrm{mol}\;\mathrm{s}^{-1}$
Lifetime	3.7	0.28	hours
Background	1.6	0.09	$\cdot 10^{22} \text{ molecules cm}^{-1}$

This  $R^2$  value indicates that the column model can replicate 55% of the variability in the observed TROPOMI line density. It is clear from Figure 4.3 that neither the increase over the source area nor the downwind decrease follow the exponential functions from the column model

completely. The TROPOMI line density shows the steepest increase at x=25 km, in the middle of the source area. Hence the TROPOMI line density is lower than the simulated line density over the first part of the source area, then becomes higher until it peaks at x=40 km and finally drops again below the observed line density for the last 10 km of the source area, suggesting that emissions are not constant over the city. This fluctuation in the TROPOMI line density is also present in the downwind decay part of the plume and not reproduced by the column model.

The emission estimate for 22 November 2017 from the TROPOMI line density is based on the average  $NO_x/NO_2$  ratio of 1.32. However, the observations at the Eiffel Tower at 13.00 UTC give a ratio of 1.72, which would result in a 30% higher emission (99.3 mol s<sup>-1</sup>).

#### 4.3 Discussion

#### 4.3.1 Quantifying $NO_x$ emissions from Paris

The  $NO_x$  emission estimate from the TROPOMI line density from 22 November 2017 is 38% higher than the average EDGAR emission of 54.1 mol s<sup>-1</sup> over the same area. However, it should be noted that these two emission numbers require a different interpretation. On the one hand, the emission estimate from the TROPOMI observations represents the emissions during the hours before the time of the measurement around noon. So for 22 November 2017, this would approximately be the period from 11:00 to 12:52 ( $x_s/U = 1.9$  hours). The EDGAR emissions, on the other hand, are an annual average, even though they are not constant over time in reality. Van der Gon et al. (2006) provides generalised temporal emission profiles for European air quality modelling, showing that  $NO_x$  emissions are above the average during winter, on weekdays and during daytime. Based on these temporal factors, the emission on a weekday in November between 12.00 and 13.00 hour is 30% higher then the annual average, where this variability is 5% monthly, 6% weekly and 19% diurnal. Hence the top-down emission estimate of 74.6 mol s<sup>-1</sup> from the TROPOMI observations would be slightly higher than the emission from Paris in the EDGAR inventory, which would be 70.3 mol s<sup>-1</sup>, when accounting for temporal variability. When temporal variability is considered, the TROPOMI  $NO_x$  estimate is hence consistent with EDGAR within 10%.

Yet the TROPOMI emission estimate might also be too low, since the  $NO_x/NO_2$  ratio as observed at the Eiffel Tower is higher than the average ratio that was applied. When the Eiffel Tower  $NO_x/NO_2$  ratio is used (1.72 instead of 1.32), the top-down emission and lifetime estimate would also be 30% higher (97.3 mol s<sup>-1</sup> and 4.9 hours). A study by Shaiganfar et al. (2017) with the CHIMERE model also found higher ratios during winter (1.51) than during summer (1.32) for the Paris region, which can be explained by higher ozone mixing ratios in summer. In addition, they found that the daily values could deviate by up to 30% from the seasonal mean. Thus, this top-down emission monitoring approach might require a more refined  $NO_x/NO_2$  rationing method. This ratio could for instance be obtained from tall-tower observations such as the Eiffel Tower or an atmospheric chemistry model, although the accuracy of such methods needs to be evaluated.

The strongest increase in the TROPOMI line density over the city (at x=25 km) is approximately at the centre of Paris, indicating that emissions are likely stronger at the centre than at the edges of the city. Such a spatial variability in the NO<sub>x</sub> emissions is indeed suggested in the EDGAR inventory, as shown in Figure 3.2. However, the column model considers the source area as a single homogeneous emission source. In chapter 6 an adapted column model will be applied to account for the effect of spatial emission variability within the city, which has the potential to improve the ability of the model to reproduce the TROPOMI line density.

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#### 4.3.2 Estimating the $NO_x$ lifetime

The estimated  $NO_x$  lifetime of 3.7 hours for this day in November is relatively short compared to other studies, which found mean  $NO_x$  lifetimes of 8.5 (Beirle et al., 2004), 8 (Beirle et al., 2011) and 13.1 hours (Schaub et al., 2007) for wintertime conditions. The case with typical urban conditions at noon of Seinfeld and Pandis (2006), from which the average  $NO_x/NO_2$  is derived, even gives a lifetime of 1.8 days. The short lifetime estimate from the TROPOMI line density could be explained by factors that influence the decrease of the line density beyond the source area. The  $NO_2$  molecules at the downwind end of the line density are emitted hours earlier (for x=200 km at approximately 8:30 hour) when  $NO_x$  has a longer lifetime. Indeed, when only equation 4.3 is fitted for  $x > x_s$ , then the lifetime becomes  $8.3 \pm 2.07$  hours. This raises the question whether the  $NO_x$  lifetime is accurately described by equation 4.4 and 4.5.

From Figure 4.2 it is also clear that the column model does not capture all fluctuations in the TROPOMI line density. Diurnal variability in emissions and lifetime are not incorporated in the model, while the emissions are higher (Van der Gon et al., 2006) and the lifetime is longer in the morning (Boersma et al., 2008). A longer lifetime of  $NO_x$  in the morning is due to the dependency of its primary sink (reaction 4.2) on the photo-chemical formation of the OH radical (Jacob, 1999). This diurnal cycle in emissions and lifetime is anticipated to lead to a higher line density of  $NO_2$  molecules at the end of the plume compared to a situation with constant emissions and lifetime.

Furthermore, the column model does not account for emissions downwind of the source underneath the passing plume. For instance, the city of Amiens or the A1 highway that runs in the same direction as the line density also add  $NO_x$  to the plume. Such downwind emissions that continuously add  $NO_x$  to the downwind plume would cause an apparent slower decay, or potentially even local increases, in the line density. Similar to the effect of the diurnal variation in emissions and lifetime, this would lead to a higher estimated  $NO_x$  lifetime.

Nevertheless, a different interpretation of the background line density, which is similar to  $L_{NO_2}(0)$  in the column model, could have an inverse effect. This background line density is supposed to be constant over the city and along the plume, which would be a valid assumption when it would fully represent the atmospheric background. The number of  $NO_2$  molecules in the tropospheric column above the boundary layer could for instance represent this atmospheric background, because of a longer free tropospheric lifetime due to lower OH concentrations and temperatures compared to the boundary layer. However, figure 4.1 shows that the measured  $NO_2$  columns at x=0 are higher than the minimum further away from Paris. Indeed, the average line density upwind of the source area of Paris for a length of 100 km is  $1.2 \pm 0.3 \cdot 10^{22} \text{ NO}_2$  molecules cm<sup>-1</sup>. This indicates that L(0) should not solely be regarded as a constant background, but also as an initial line density from upwind emission sources that enters the city and is immediately subject to chemical decay. When the  $NO_x$  loss rate k would be applied to the B term in the column model, then instead of all, only part of the downwind exponential decay would be explained by the first term on the right side of equation 4.5, resulting in a shorter top-down estimated lifetime.

Hence the relatively low  $NO_x$  lifetime that was estimated here, is influenced by four effects on the downwind decay that were not accounted for in the column model:

- The diurnal emission cycle
- Variability in lifetime
- Downwind emissions
- Background decay

The first three effects are expected to result in an underestimation of the lifetime and hence an overestimation of the emission, whereas the fourth is anticipated to shown an opposite effect. These influences on the downwind decay will be investigated further in chapter 5.

# 5 Influences on downwind decay

The previous chapter showed how the emission and lifetime can be estimated from the TROPOMI NO<sub>2</sub> line density. Here the sensitivity of such an emission and effective lifetime estimate to different influences on the downwind decay is tested. Four bottom-up line densities are calculated based on the EDGAR emissions inventory, from which the lifetime and emission are derived with the top-down column model. The reference case uses the total emission for the source area and a predefined lifetime to simulate a line density. The annual average emissions from the EDGAR inventory are multiplied by a factor of 1.3 in order to be representative for a weekday in November around noon (Van der Gon et al., 2011). Four experiments are performed where this reference case is adapted with:

- 1. Downwind emissions: adding EDGAR's spatial emission pattern and downwind emissions (on average 10% relative to Paris)
- 2. Diurnal emissions: including 20% higher emissions during rush hour
- 3. Variability in lifetime: accounting for a 8.3 times lower OH concentration at sunrise than at noon
- 4. Background decay: applying chemical decay to the background line density

The original column model is then applied in the inverse mode to determine the emission and lifetime from these four simulated line densities. These results are compared to the reference case to determine the sensitivity of the emission and lifetime to these four different influences.

# 5.1 Methodology

#### 5.1.1 Downwind emissions

The column model assumes a constant emission strength over the source area and that no downwind emission sources are present. To test the sensitivity of emission and lifetime estimates from the top-down column model to this assumption, a bottom-up line density is simulated based on the EDGAR 2012 emission inventory including the emissions downwind of Paris and its spatially variation. This line density is simulated with an adapted form of the column model, where a background line density of  $1.7 \cdot 10^{22}$  NO<sub>2</sub> molecules cm<sup>-1</sup> and a lifetime of 9.8 hours is assumed.

The lifetime of 9.8 hours is calculated with equation 4.3, with an OH concentration of 1.5  $\cdot 10^6$  molecules cm<sup>-3</sup> (Kanaya et al., 2007), rate constant of 2.5  $\cdot 10^{-11}$  cm<sup>3</sup> molecules<sup>-1</sup> s<sup>-1</sup> for reaction 4.2 (for surface conditions, Seinfeld and Pandis, 2006) and NO<sub>x</sub>/NO<sub>2</sub> ratio of 1.32 (Beirle et al., 2011). This OH concentration is the daytime maximum as observed during the IMPACT IV campaign in central Tokyo from January to February 2004 (Kanaya et al., 2007), which is assumed to be representative for a megacity during winter. Although no OH measurements in Paris for November were available, observations in Paris during the MEGAPOLI campaign in July 2009 (Michoud et al., 2012) show similar values as observations in Tokyo from July to August 2004 (Kanaya et al., 2007), respectively 3.5–10.6 and 6.3  $\cdot 10^6$  molecules cm<sup>-3</sup>.

The emissions are taken from an emission line density, which integrates the emissions from the EDGAR inventory perpendicular to the wind direction (y) for every x along the wind direction through the centre of Paris, similar to equation 4.1. The emission line density starts 30 km from the centre of Paris and has a total length of 200 km, corresponding the the domain of the TROPOMI

5.1. Methodology

line density in the previous chapter. An emission line density is calculated for each interval  $(\delta x)$  of 7 km, similar to the resolution of the EDGAR inventory of 0.1 by 0.1 degrees (approximately 7 by 11 km). Since the EDGAR emissions are reported as mass of NO<sub>x</sub>, these numbers are converted to NO<sub>x</sub> molecules with the molecular weight of NO<sub>2</sub>, as specified in the EDGAR data set. The resulting emission line density is shown in figure 5.1, showing a clear emission peak over Paris and 10 times lower downwind emissions. The exact origin of the peak around x = 175 km could not be found, since there is not city or other large emission source at this location (50.15N 2.65E in Figure 3.2).

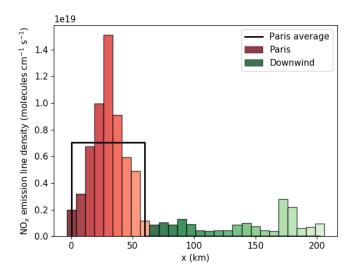


Figure 5.1: Emission line density from the EDGAR  $NO_x$  emission inventory along the wind direction (x) from 2012 and the average EDGAR  $NO_x$  emission for Paris and zero downwind emissions.

For all n emissions  $(E_i)$  at  $x_i$  on the emission line density, a NO<sub>2</sub> line density is simulated, which is shown in Figure 5.2. In contrast to Jacob's column model (1999), which has a source area from 0 to  $x_s$ , in this approach the source area runs from  $x_i - \frac{\delta x}{2}$  to  $x_i + \frac{\delta x}{2}$  for each emission on the emission line density, which correspond to the approximate grid size of the EDGAR inventory  $(\delta x = 7 \text{ km})$ . The line density  $L_i$  for  $E_i$  as a function of x is then given by the following three formulas:

$$L_i(x, E_i) = 0 \text{ for } 0 \le x < x_i - \frac{\delta x}{2}$$
 (5.1)

$$L_i(x, E_i) = \frac{E_i}{k} \left( 1 - e^{-\frac{k(x - x_i + \frac{\delta x}{2})}{U}} \right) \text{ for } x_i - \frac{\delta x}{2} \le x \le x_i + \frac{\delta x}{2}$$
 (5.2)

$$L_i(x, E_i) = L_i(x_i, E_i)e^{-\frac{k(x - x_i - \frac{\delta x}{2})}{U}} \text{ for } x > x_i + \frac{\delta x}{2}$$
 (5.3)

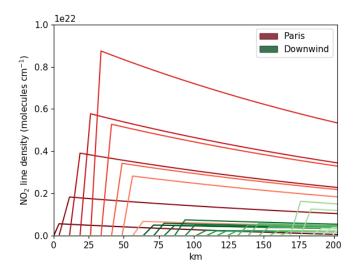


Figure 5.2: All n simulated NO<sub>2</sub> line densities for each  $E_i$  on the emission line density, where red depicts the line densities for emission from Paris ( $x \le x_s$ ) and green the line densities for downwind emissions ( $x > x_s$ ).

All the resulting simulated line densities for each  $E_i$  are then summed to obtain the total bottom-up line density:

$$L(x) = \sum_{i=1}^{n} L_i(x, E_i) + B$$
 (5.4)

Finally, the inverse method is used to obtain an estimate for the emission and lifetime by fitting the column model functions to the simulated bottom-up line density, which are compared with the initially assumed lifetime of 9.8 hours and the temporally corrected spatially averaged emission over Paris (70.3 mol  $NO_x$  s<sup>-1</sup>, from chapter 4). An uncertainty of 50.7% in the simulated line density is used, based on the uncertainty in the EDGAR emission inventory (Crippa et al., 2018).

#### 5.1.2 Diurnal emissions

The column model implicitly assumes that the emission source is continuous. In reality, Paris'  $NO_x$  emissions follow a diurnal cycle, with strongest emissions during rush hour. This can be seen in the observed  $NO_2$  concentrations at the top of the Eiffel Tower, as shown in Figure 5.3 for 22 November 2017, where the concentrations during rush hour are approximately 20% higher than around noon. Similar to the Eiffel Tower observations, the temporal emission pattern from the TNO report by van der Gon et al. (2011) gives 18% higher rush-hour emissions (7.00 to 9.00 UTC) compared to noon (Figure 5.3). In this third experiment, this pattern is used to scale the EDGAR emissions in the column model:

$$E(t) = E_{\text{EDGAR}}(13h) \cdot \frac{E_{\text{TNO}}(t)}{E_{\text{TNO}}(13h)} \text{ with } t(x) = \frac{x - x_s}{U}$$
(5.5)

From the resulting new line density, the emission and lifetime of NO<sub>2</sub> is derived with the inverse column model to assess the influence of temporal emission variability on those two parameters.

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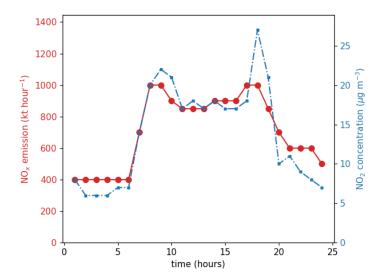


Figure 5.3: Diurnal cycle in  $NO_x$  emissions from the EDGAR inventory with temporal correction and  $NO_2$  concentration observed on 300 m at the Eiffel Tower in Paris on 22 November 2017.

#### 5.1.3 Variability in lifetime

Similar to the previous experiment on the diurnal emission cycle, also the lifetime varies throughout the day. During daytime, the  $NO_x$  lifetime is determined by the oxidation with OH to  $HNO_3$ , as described by reaction 4.2. Since the formation of OH requires solar radiation,  $\tau_{NO_x}$  is shortest around noon and longest at dawn. Similar to  $NO_x$  lifetime derivation in paragraph 5.1.1 and because no OH observations in Paris on 22 November 2017 were available, figures from the IM-PACT campaign in Tokyo are used that gives an OH concentration of 0.18 ·10<sup>6</sup> molecules cm<sup>-3</sup> during nighttime (Kanaya et al., 2007). This is 8.3 times lower than the maximum daytime OH concentration. Although it is likely that the OH concentration in Paris on 22 November 2017 would be different from these observations over Tokyo from January to February 2004, the diurnal variation should give a good representation of the variation in lifetime, and hence the sensitivity of the estimated column model parameters.

Since the TROPOMI observations consist mainly of  $NO_x$  emitted in the hours before its early-afternoon overpass time, the linear change in OH concentration per unit of time  $(\frac{\delta[OH]}{\delta t})$  is given by the difference between the OH concentration at noon (12.36 UTC) and at dawn (8.10 UTC) over the difference in time. Before dawn and after noon the OH concentration is kept constant. Based on equation 4.3 the variable rate constant  $k_v$  is implemented in the model with the following formula as a function of dt, which is the time before the TROPOMI overpass time in seconds:

$$k_v(dt) = \frac{k_2([OH]_{\text{noon}} - \frac{\delta[OH]}{\delta t}dt)}{[NO_x]/[NO_2]}$$
(5.6)

This results in a  $NO_x$  lifetime of 81.5 hours at dawn and 9.8 hours at noon, shown in Figure 6.2. This figure runs from the TROPOMI overpass time at 12.52 back to 7.05 UTC, since part of the end of the plume (x = 200 km) is emitted 5.8 hours earlier (x/U).

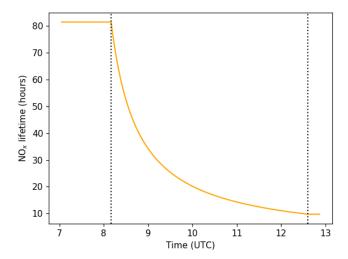


Figure 5.4: Variability in  $NO_x$  lifetime for wintertime conditions, decreasing from sunrise (8.10 UTC, dotted line) to solar noon (12.36 UTC, dotted line) from the morning until the TROPOMI overpass time.

Because the column model by Jacob (1999) considers the chemical decay to be constant, the lifetime as a function of t is implemented into the derivatives of equation 4.4 and 4.5. These derivatives describe the change in the line density per distance along the wind direction  $(\frac{dL}{dx})$ . This change is determined by the emission (E) for  $x \leq x_s$  and the chemical decay  $(k \cdot L)$  as a function of  $x_i$  and x. The variable  $x_i$  denotes the location of the derivative of L, whereas x defines the location of L once it is observed by TROPOMI, so that  $\frac{x_i - x}{U}$  is dt, the time that it would take for the line density at  $x_i$  to reach x. In other words, each line density as observed at 12.52 UTC would at an earlier time (12.52 UTC - dt) be located at  $x_i$  where it would shrink because of chemical decay and, if located over the source area, grow because of emissions. This is described by the following two equations, with  $dt = \frac{x_i - x}{U}$ :

$$\frac{dL}{dx}(x_i, x) = \frac{E - k_v(dt) \cdot L(x_{i-1})}{U} \text{ for } 0 \le x \le x_s$$
(5.7)

$$\frac{dL}{dx}(x_i, x) = \frac{-k_v(dt) \cdot L(x_{i-1})}{U} \text{ for } x > x_s$$
 (5.8)

The line density (L) as a function of x at dt = 0 is then given as the sum of each derivative at  $x_i$  up to  $x_n = x$ :

$$L(x) = \sum_{i=1}^{n} \frac{dL}{dx} (x_i, x) \cdot (x_i - x_{i-1}) \text{ where } 0 \le x_i \le x$$
 (5.9)

Similar to the previous two experiments, the simple column model is inversely applied to obtain estimates for the emission, lifetime and background line density. These figures are subsequently compared to the reference case and the other experiments to determine the sensitivity of these estimated parameters to lifetime variability.

#### 5.1.4 Background decay

As the emission and lifetime term in the column model are evaluated in the previous three experiments, this last and fourth experiment focuses on the background component (B). From the TROPOMI line density from 22 November 2017, the column model estimates the line density to

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be  $1.6 \cdot 10^{22} \text{ NO}_2$  molecules cm<sup>-1</sup> (Table 4.1). However, by looking at both the TROPOMI and fitted column model functions, it appears that the background line density is mainly determined by the initial line density at x = 0 km. This raises the question whether this value truly represents the background.

The variability in the upwind line density could be explained by a different regional background level, but also partly by upwind emission sources. When the background line density would originate completely from upwind emissions, then it should rather be seen as an 'initial' line density, which chemically decays as it moves along with the wind. This interpretation is implemented in the model as follows:

$$L_{NO_x}(x) = \frac{E}{k} \left( 1 - e^{-\frac{kx}{U}} \right) + Be^{\frac{-kx}{U}} \text{ for } 0 \le x \le x_s$$
 (5.10)

$$L_{NO_x}(x) = L_{NO_x}(x_s)e^{-\frac{k(x-x_s)}{U}} + Be^{\frac{-kx}{U}} \text{ for } x > x_s$$
 (5.11)

Although it seems probable that at least part of the background line density of  $1.7 \cdot 10^{22} \text{ NO}_2$  molecules cm<sup>-1</sup> originates from upwind emission sources, this part would be difficult to distinguish from the component that is uniform in space. This spatially uniform part could be the result from an error in the partitioning between the tropospheric and stratospheric column.

In this experiment, the full background component is decaying as it moves in the x direction. The resulting simulated density in combination with the column model is used to estimate the three parameters E, k and B, to test how sensitive this approach is to the assumption that the background component remains constant.

#### 5.2 Results

When downwind emissions are added to the column model, the simulated line density for experiment 1 shows a weaker decay downwind of the source area (Figure 5.5a). At the end of the plume, almost one third of the line density originates from downwind emissions. When the original column model is fitted to the line density with downwind emissions, the resulting NO<sub>x</sub> lifetime is longer. Hence neglecting downwind emissions leads to an overestimation of the lifetime. Interestingly, the higher lifetime in this experiment does not lead to a lower emission. While E is 3.7 mol s<sup>-1</sup> higher than the reference case (Table 5.1), the opposite would be expected, since a similar change in magnitude requires a larger E when k becomes smaller ( $\frac{dL}{dt} = E - kL$ ). Yet it appears that this effect is compensated with the lower estimated background line density by the fitted column model. This lower value for B leads to a better fit over the source area, since the simulated line density has an S-shape, with the strongest increase over the middle of the city.

Similar to the result of adding downwind emissions, experiment 2 with a diurnal emission cycle shows a higher lifetime. Because the effect of the spatial distribution in emissions over the source area does not play a role here, the  $NO_x$  emission is estimated to be 4.5 mol s<sup>-1</sup> smaller than the reference case (Table 5.1). The simulated line density, shown in Figure 5.5b, is higher at the end of the plume, because this part originates from rush-hour  $NO_x$  emissions. This results in a lifetime that is 6.2 hours longer.

Accounting in the model for a variable lifetime in experiment 3 gives a quite similar effect as the diurnal emission cycle. Since earlier emissions have a longer lifetime, the end of the simulated  $NO_2$  line density becomes larger (Figure 5.5c). The lifetime that is obtained from applying the inverse column model is 7.1 hours longer, and the emission 3.7 mol s<sup>-1</sup> smaller compared to the reference case (Table 5.1).

In contrast to the previous three experiments, experiment 4 gives a lower line density because a decay factor is applied to the background parameter (Figure 5.5d). The resulting estimated

lifetime is 4.3 hours shorter and the emission 1.1 mol  $\rm s^{-1}$  smaller than the reference case (Table 5.1).

Table 5.1: Column model estimates and standard deviation of  $NO_x$  emission, lifetime and background line density fitted to the line densities of the four experiments, in comparison to the reference case.

Experiment	Emission	Lifetime	Background
	$(\text{mol s}^{-1})$	(hours)	$(\cdot 10^{22} \text{ molecules cm}^{-1})$
Reference	$70.3 \pm 11.35$	$9.8 \pm 4.27$	$1.7 \pm 0.22$
1. Downwind emissions	$74.0 \pm 10.59$	$29.3 \pm 31.14$	$1.5 \pm 0.22$
2. Diurnal emissions	$65.8 \pm 22.37$	$16.0 \pm 22.22$	$1.7 \pm 0.48$
3. Variable lifetime	$66.6 \pm 22.40$	$16.9 \pm 24.44$	$1.7 \pm 0.48$
4. Background decay	$69.1 \pm 12.16$	$5.5 \pm 1.69$	$1.7 \pm 0.22$

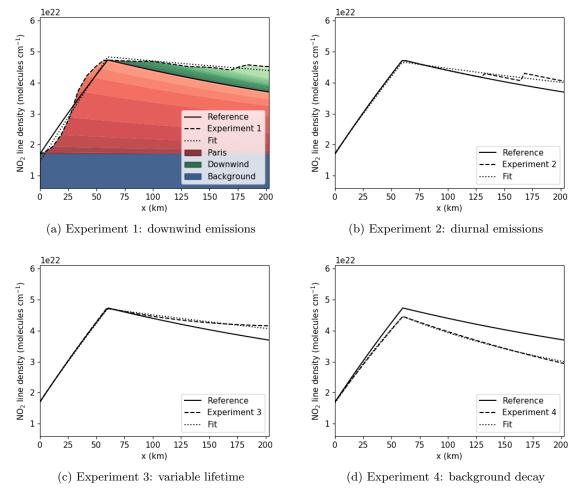


Figure 5.5: The simulated line densities of the reference case, the four experiments and the inverse column model that is fitted to obtain estimates for the emission, lifetime and background line density.

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#### 5.3 Discussion

The four different experiments show how the estimated  $NO_x$  emission, lifetime and  $NO_2$  background line density are influenced by spatial and temporal variability in emissions, addition of emissions to the downwind line density and background decay. Accounting for downwind emissions in experiment 1 leads to the largest difference in lifetime compared to the reference case, namely 3.0 times longer. Although one would expected that this longer lifetime results in a lower emission based on the column model equations, experiment 1 gives a higher emission than the reference case because the fit adapts to the spatial emission pattern by lowering the background line density. When this effect is excluded by keeping B constant in the fit, then an emission of 65.3  $\pm$  7.08 mol s<sup>-1</sup> is estimated from the simulated line density with downwind emissions. Hence even a very large overestimation of the lifetime by a factor of 3.0 due to downwind emissions results in a relatively small underestimation of the emission by 7.1%.

This finding that a relatively large overestimation in lifetime results in a slight underestimation in emission is confirmed in experiment 2 and 3, which give respectively a 6.4% and 5.3% lower emission compared to the reference case. In comparison, the study by Beirle et al. (2011), which estimated  $NO_x$  emissions and lifetimes for multiple megacities from line densities based on averaged satellite observations for multiple years, also concludes that "possible different nighttime behaviour of  $\tau$  and E has almost no impact" on their estimated parameters (supporting online material, p. 6). Similar to experiment 2, their simulation with 10 times lower nighttime emissions leads to approximately 10% higher emissions, and similar to experiment 3, their estimated emission is approximately 10% lower from a simulated line density with a 10 times higher nocturnal lifetime (Beirle et al. 2011).

Beirle et al. (2011) also found that the lifetimes obtained from the fits to the simulated line density with longer lifetimes and lower emissions during the night deviate less than two hours from their default lifetime of 5 hours. Although the differences in lifetime between experiment 2 and 3 with the reference case are much larger than Beirle et al. (2011), this can be explained by the shorter initial lifetime of 5 hours compared to 9.8 hours. When experiment 2 and 3 are performed again, but with a doubled OH concentration that corresponds to a lifetime of 4.8 hours, the difference in lifetime is in the same order of magnitude, namely 1.0 and 3.5 hours longer.

However, the large differences in  $NO_x$  lifetime estimates between the four experiments and the reference case show that it remains difficult to get an accurate estimate with fitting an exponential function through the downwind decay in a line density. The experiments presented here could be combined to improve the column model and potentially enable a more accurate lifetime estimation.

Besides improving the model to obtain a better lifetime estimate from the downwind line density, it might also be an option to obtain an OH concentration from other sources, such as the CAMS model, to calculate the  $NO_x$  lifetime. When the lifetime is not estimated from the line density, the emission can be estimated solely from the line density over the source area. Downwind emissions, the diurnal emission cycle and variability in lifetime would then not influence the estimated emission with the column model anymore. An effect that should be considered is the spatial variability in emissions over the city, as that has shown to have a large influence in experiment 1. This approach will be evaluated in the next chapter, where an emission will be estimated with the column model just from the increase in line density over Paris.

# 6 Accumulation over the source area

In chapter 4, it is shown how the column model as described by Jacob (1999) can be used to derive the  $NO_x$  emission and lifetime from a line density of TROPOMI observations. Because chapter 5 proved the large effects in the downwind component of the line density on the column model lifetime estimates, this sixth chapter evaluates the accumulation of  $NO_2$  just over Paris in the wind direction.

Besides the TROPOMI observations of 22 November 2017, two longer periods from 22 to 26 February and from 17 to 24 April 2018 are evaluated. In these periods three days were excluded, as described in paragraph 3.5. Firstly, the TROPOMI observations are validated by comparing the surface  $NO_2$  measurements in Paris with the line density for each of the 11 days to test whether or not the built-up of  $NO_2$  in the wind direction can also be seen in the surface observations. Secondly, the emission is estimated just from the TROPOMI line density over the source area with a column model as well as an adapted column model that incorporates the spatial emission pattern within the city, both with a simulated  $NO_x$  lifetime from a chemical transport model.

This new approach, where the the downwind decay component is disregarded and the  $NO_x$  emission is estimated just from the accumulation over the source area, would be an optimal use of increased resolution of TROPOMI compared to previous instrument. Indeed, from Figure 6.2 one can see that the source length  $x_s$  of Paris encompasses eight TROPOMI observations, which is also reflected by the strongest increase in the line density over the middle of the city.

# 6.1 Methodology

#### 6.1.1 Comparison with surface observations

To compare the surface observations with the TROPOMI line density, the distance x along the wind direction is calculated for each observatory. The resulting observed NO<sub>2</sub> concentrations as a function of x are then averaged for each interval on the TROPOMI line density, if the difference in distance x between the surface observation and TROPOMI line density is smaller than 3.5 km (half of the integration interval of the TROPOMI line density). This average is then compared with the TROPOMI line density, to validate that this pattern is also visible in the measured surface concentrations.

#### 6.1.2 Estimating emission from source area accumulation

In order to test whether the  $NO_x$  emission can be estimated just from the accumulation over Paris based on the TROPOMI line density, the column model is applied, similar to the method described in chapter 4 but now focusing on the source area. In other words, just equation 4.4 of the column model is applied.

Since the lifetime cannot be estimated from the decline downwind of the source area, predefined lifetimes are assumed (Table 6.1). The  $NO_x$  lifetime for 22 November 2017 is 9.8 hours, as used in chapter 4 and 5. The lifetimes for the days in February and April 2018 are taken from Lorente et al. (unpublished paper, 2018), which is based on Equation 4.4, the mean boundary layer concentration of  $NO_x$ ,  $NO_2$  and OH simulated by the CAMS model over Paris at noon and a rate constant  $(k_2)$  of  $2.6 \cdot 10^{-11}$  cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>. Because 19 April was not included in this study, the average lifetime from the other days in April was taken here. The assumed lifetimes show a strong difference

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between February and April because of more solar radiation that is required for the formation of OH.

It should also be noted that the sounding observations of wind speed and direction at Trappes (Table 3.1), which are used to derive the line densities and estimate emissions with the column model, were unavailable for 22 February, 23 February and 22 April. Wind date for these data is supplemented by Lorente et al. (unpublished paper, 2018), whom calculated the mean wind for the boundary layer at 12.00 UTC from the 6-hourly ECMWF-Interim reanalysis.

Table 6.1: Predefined  $NO_x$  lifetimes used to estimate emission with the column model from the line density over the source area.

Month	November 2017	February 2018		April 2018							
Day	22	22	23	25	26	17	18	19	21	22	24
$\overline{NO_x}$ lifetime (hours)	9.8	16.0	10.9	11.4	10.3	1.6	2.0	2.6	2.9	2.7	2.8

#### 6.1.3 Spatial column model for emission pattern over the source area

The TROPOMI line density of 22 November 2017, illustrated in Figure 4.3 from chapter 4, showed the strongest increase over the centre of Paris. Indeed, also the bottom-up simulated line density of experiment 1 in chapter 5, which included the spatial distribution in emissions from the EDGAR inventory, showed an S-shape with the strongest increase of  $NO_2$  in the x direction over the centre of the source area. This hints that including the spatial emission pattern in the column model would result in a better fit with the TROPOMI line density, and hence a more accurate emission estimate.

To test this assumption, and adapted form of the column model accounting for the spatial emission variability is used to estimate emissions. A large number of line densities is simulated for a large number of different total emission strengths, of which the estimate is the emission corresponding to the simulated line density with the lowest  $\chi^2$  (Equation 4.6) compared to the TROPOMI line density.

The line densities are simulated similarly to the method described in section 4.1.1, where the simulated line density is the sum of the line densities for each emission at distance x on the emission line density. This is described by Equation 5.11. The emission line density is based on the EDGAR inventory, and averaged for all wind directions with intervals of 10 degrees (Figure 6.1).

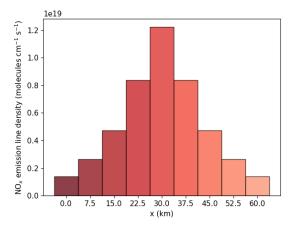


Figure 6.1:  $NO_x$  emission line densities averaged for all directions.

#### 6.2 Results

The tropospheric NO<sub>2</sub> columns measured by TROPOMI show an increase in the wind direction, as illustrated in Figure 6.2 and 6.3, for all days that were analysed, except 22 April 2018. On this day it is difficult to notice a trend over the domain of the city, although higher columns can be seen downwind (northeast) of the city. These figures also show that the wind direction that is used corresponds well with the direction of the NO<sub>2</sub> plume that is observed with TROPOMI; for most days both point in the same direction. For two days, however, one can note a slight deviation between the wind's and plume's direction. On 22 February 2017 the plume is pointed more towards the south and on 21 April 2018 the direction of the plume is tilted towards the north, in comparison with the direction of the wind. Yet this does not appear to be a problem for the line density derivation, since the full width of the plume is still covered within the domain of the line density at the end of the source area (x=60 km).

The surface NO<sub>2</sub> measurements show a similar pattern as TROPOMI's NO<sub>2</sub> columns: higher values closer to the centre and at the downwind side of the city. This is shown in Figure 6.3 for all days in February and April 2018, and in Figure 6.2 for 22 November 2017, since this day has higher surface NO<sub>2</sub> concentrations compared to the other days and therefore requires a different scaling. When comparing all days in Figure 6.3, it becomes clear that also the magnitude of measured  $NO_2$ correspond between the surface concentrations and tropospheric columns. For instance 22 April 2018, when both TROPOMI and surface measurements give low values. In addition, for most days also the measured surface concentrations averaged at distance x show an increase similar to the TROPOMI line densities (Figure 6.5). However, this is not the case for 19 April 2018, perhaps due to the comma-shape of the downwind plume (Figure 6.3) that is caused by turning wind (from 160 at 8.00 UTC to 220 at 13.00 UTC at the surface station in Trappes), 21 April 2018, where one single station around x=47 km lowers the average value at the downwind side of the city, and 24 April 2018, which hints at an error in the TROPOMI observations that are unexpectedly low over the centre of Paris. Nonetheless, it can be concluded that overall the increase in NO<sub>2</sub> abundance over the city of Paris in the wind direction is not only observed by TROPOMI, but also by surface observations.

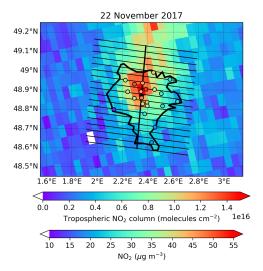


Figure 6.2: Measured tropsopheric  $NO_2$  column by TROPOMI and surface observations on 22 November 2017. The thick line denotes the wind direction (x) and the thin lines with intervals of 3.5 km represent each line density for which all pixels perpendicular to the x direction are integrated.

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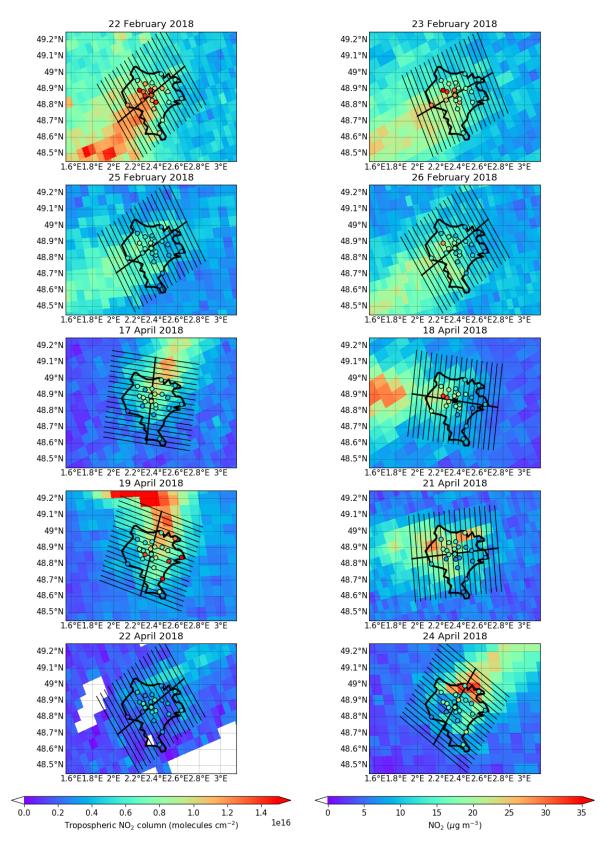


Figure 6.3: Measured tropsopheric  $NO_2$  column by TROPOMI and surface observations for the two periods in February and April 2018. The thick line denotes the wind direction (x) and the thin lines with intervals of 3.5 km represent each line density for which all pixels perpendicular to the x direction are integrated.

Since the surface observations validate the trend in the TROPOMI line density, the next step is to estimate the  $NO_x$  emissions from the TROPOMI line density with both the column model and spatial column model, the latter accounting for the spatial emission pattern. After fitting the column model functions to the TROPOMI line density (Figure 6.5), the resulting  $NO_x$  emission estimates from the column model are higher or almost the same as the spatial column model for most days (Table 6.2). For about half of the analysed days, namely all days in February 2018 and both 17 and 19 April 2018, the spatial column model has lower  $\chi^2$  values, indicating a better fit to the TROPOMI line density than the column model without the spatial emission distribution. Indeed, for these days the TROPOMI line densities show a clear symmetrical S-shape, while the days where the spatial column model performs worse show some decline at the downwind side of the city. Thus, in contrast to the hypothesis based on the discussions of the results from chapter 4 and 5, it cannot be concluded that including spatial variation of emissions from the EDGAR inventory within the source area in the column model, while using a predefined lifetime, leads to an improvement.

When interpreting the emission estimates from Table 6.2, it is good to note that ERA-Interim wind data was used for 22 and 23 February 2018, as well as 22 April 2018. The wind speeds for these days appears to be higher, in comparison to the sounding observations and surface wind measurements (Table 3.1). According to the column model equations, a higher wind speed would result in a lower emission. This could partly explain that 22 April 2018 is the lowest estimated emission. In contrast, the two days in February are actually at the higher end of the emission estimates.

Table 6.2: Emission estimates for all days with the column model and spatial column model and the respective  $\chi^2$  value for the model fit with the TROPOMI line density.

Date Day		Column model		Spatial column model		
		Emission (mol $s^{-1}$ )	$\chi^2$	Emission (mol $s^{-1}$ )	$\chi^2$	
22/11/17	Wednesday	60.2	25.5	56.6	28.5	
22/02/18	Thursday	67.7	24.7	58.2	3.3	
23/02/18	Friday	59.3	18.1	51.8	4.7	
25/02/18	Sunday	31.2	11.2	30.1	7.5	
26/02/18	Monday	51.5	30.7	43.2	17.5	
17/04/18	Tuesday	59.3	44.2	54.6	6.1	
18/04/18	Wednesday	68.1	18.1	68.6	33.0	
19/04/18	Thursday	50.7	26.3	46.1	21.4	
21/04/18	Saturday	38.7	12.9	35.8	15.4	
22/04/18	Sunday	29.3	19.0	25.4	24.6	
24/04/18	Tuesday	48.4	15.2	51.7	43.0	

From the magnitude of the emissions for each day of the week, illustrated in Figure 6.4, it becomes clear that the  $NO_x$  emissions during the weekend are half as large as during weekdays (54% for both the column model and the spatial column model). The standard deviations of these averages are all quite low, where the highest standard deviation is 9.0 mol s<sup>-1</sup> for the column model emission estimates during weekdays.

The estimated  $NO_x$  emissions are also substantially lower then the averaged emission from the EDGAR inventory for Paris from 2012 with a correction for temporal variation of 4.7% for winter, 19.3% for noon, 6.1% for weekdays and -15.2% for weekends (van der Gon et al., 2011). This

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results in an average  $NO_x$  EDGAR emission for Paris is 73.1 mol s<sup>-1</sup> for weekdays and 61.1 mol s<sup>-1</sup> for the weekend. In comparison, the column model estimated  $NO_x$  emission is respectively 24% and 51% lower than the temporally corrected EDGAR emissions. Also the difference between weekday and weekend emissions is much larger in the average column model estimates: 2.0 times for the column model and just 1.2 times for the EDGAR inventory.

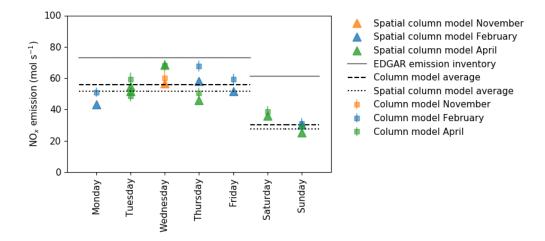


Figure 6.4: Emission estimates for Paris with the column model and spatial column model from the TROPOMI line densities for 22 November 2017 and the two periods in February and April 2018, as well as the average for weekdays and weekend in comparison to the emission from the EDGAR inventory with temporal correction.

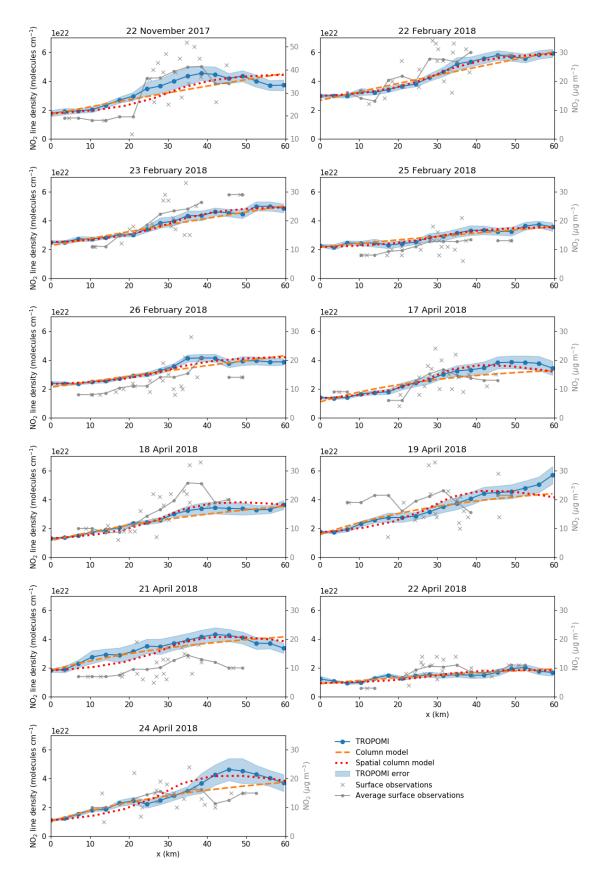


Figure 6.5:  $NO_2$  line densities over Paris from TROPOMI, measured surface concentration and column model fits.

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## 6.3 Discussion

This sixth chapter showed how the  $NO_x$  emission from Paris can be estimated just from the accumulation over the city, which is now possible because of the higher resolution of TROPOMI in comparison to its precessors. Previous studies such as by Beirle et al. (2011) required satellite  $NO_2$  observations beyond the scope of the source domain to determine the  $NO_x$  lifetime from the downwind decay. However, as shown in chapter 5, this approach adds complexity since downwind emissions and variability in time influence the estimated  $NO_x$  lifetime, processes that are not considered in the column model as introduced in chapter 4. These processes do not pose a problem when the  $NO_x$  emission is estimated from the increase in  $NO_2$  abundance over the city and the  $NO_x$  lifetime is obtained in a different way. Although the use of a lifetime from a chemical transport model brings in another uncertainty, the effect of chemical decay is only minor 1 compared to the increase due to emission. This was also confirmed in chapter 5, which concluded that even a large overestimation of the lifetime leads to only a small underestimation of the emission.

Besides the advantage that influences on the downwind plume do not perturb the emission estimate when applying the column model just to the source area, this new approach also has a second advantage. Whereas previous efforts to estimate  $NO_x$  emissions with satellites relied on averaged data, often for multiple years, to get a sufficiently fine resolution, here daily emission estimates can be made just from a single TROPOMI overpass. This makes it possible to analyse seasonal and diurnal emission patterns, which can lead to a better understanding of contributions from different emission sources and hence the effectiveness of climate change and air quality mitigation strategies.

The weekly cycle of  $NO_x$  emissions from the column model estimates for eleven analysed days showed a factor of 2.0 higher emissions during weekdays than the weekend. This is a much stronger differences than the report by van der Gon et al. (2011), which differ by a factor of 1.2. One would expect lower  $NO_x$  emissions during the weekend, as there is less traffic than during the week. The lower difference in the temporally corrected EDGAR inventory could hence hint at a too high emission factor for traffic or an underestimation of traffic numbers in the activity data.

Not only is the weekly variation stronger, also the average  $NO_x$  emissions during weekdays and weekend of the estimates from the TROPOMI line density are well below the average reported emissions of the temporally corrected EDGAR inventory (respectively 24% and 51%). This could mean that the  $NO_x$  emissions reported in the EDGAR inventory are too high. However, one should note that the EDGAR inventory from 2012 is used. The tri-annual mean  $NO_2$  concentration decreased by about 8% at background sites in Paris from 2011-2013 to 2015-2017 (Airparif, 2018). Also the European Environmental Agency reports a decrease in  $NO_x$  emissions by 15% for France from 2012 to 2016 (EEA, 2017). This emission reduction could could explain to a certain extend the lower TROPOMI emission estimates compared to the EDGAR inventory of 2012.

However, directly comparing  $NO_x$  emission figures from the top-down estimates based on TROPOMI observations with those from the bottom-up reported EDGAR emission inventory requires caution. As mentioned before, the former is representative for the emissions around midday, whereas the latter is an annual average that can be corrected with temporal correction factors. No temporal correction factors are available from the EDGAR inventory. And it is likely that the factors used in this study from a TNO report (van der Gon et al., 2011) have a high uncertainty embedded. For instance, the diurnal cycle in emissions from road transport is based on traffic intensity time series from 1985 to 1998 in the Netherlands. It seems plausible that not only traffic intensity and emission abatement in cars decreased in the past decades, but also that the temporal variation in Paris is different than in the Netherlands.

<sup>&</sup>lt;sup>1</sup>The chemical decay rate is 7% of the emission according to  $\frac{kB}{E}$ , with k and B based on the lifetime and background used in chapter 5 and E as the average EDGAR emission over Paris from figure 5.1.

Furthermore, not only the limitations of the temporal correction factors require attention, but also the TROPOMI  $\mathrm{NO}_x$  estimates. Firstly, the accuracy of the tropospheric  $\mathrm{NO}_2$  columns as observed by TROPOMI needs to be considered. Although the validation with the surface  $\mathrm{NO}_2$  measurements in Paris aligns well with both the variation in space and magnitude of the TROPOMI observations, a further validation would be recommended, for instance with ground-based MAX-DOAS measurements.

Secondly, the wind speed is an important parameter in the column model and has a large influence on the estimated  $NO_x$  emission. When wind speeds are halved, which corresponds on some days to the difference between the wind speed at the surface and middle of the boundary layer (Table 3.1), the column model  $NO_x$  emission estimates are on average 40% lower then the values presented in Table 6.2. In order to determine a more accurate wind speed for the column model, it would be useful to have a better understanding of the vertical profile of  $NO_2$ , since wind speed increases with altitude.

Thirdly, the chemical decay in the column model needs to be considered. A disadvantage of applying the column model over the city is that estimating the  $NO_x$  lifetime from the TROPOMI line density would be difficult as the signal is minor compared to the emissions. Hence lifetimes are needed from CAMS model simulations (unpublished paper, Lorente et al., 2018), whom report an uncertainty of 50% for these lifetimes. When the lifetimes as described in Table 6.1 are lowered by 50%, the column model emission estimates are on average 21% higher then the values in Table 6.2. However, a limitations to the results from this chapter is the conversion from  $NO_2$  to  $NO_x$ . This is done with the  $NO_x/NO_2$  ratio, which is assumed to have a constant value of 1.32, while the Eiffel Tower observations give an average ratio of 1.79 for four days in February 2018 and 1.40 for five days in April 2018. Applying these higher  $NO_x/NO_2$  ratio to the column model would lead to a higher emission.

And lastly, accounting for the spatial emission pattern in the column model, with the steepest increase in line density over the centre of the source area, does only lead to a better fit (lower  $\chi^2$ ) for some days. It seems that, except for 22 November 2017 and 17 April 2018, especially on days with higher wind speeds the spatial column model performs better then the simple column model. Perhaps this is due to the fact that chemical loss becomes relatively more important at lower wind speeds.

# 7 Conclusion

The improved resolution of the new TROPOMI satellite measurements now enables us to see how the abundance of  $NO_2$  in the troposphere accumulates over a large source area, such as the city of Paris, and decays downwind. This evolution of  $NO_2$  in the direction of the wind can be described by the TROPOMI line density, from which the  $NO_x$  emission, lifetime and background can be estimated with a simple column model for a single day. After applying this method to the TROPOMI observations of 22 November 2017, the resulting estimated  $NO_x$  is just 6% higher than the EDGAR emission inventory when a temporal correction factor is applied. However, the estimated  $NO_x$  lifetime of 3.7 hours seems to be too short in comparison with literature (Beirle et al., 2011; Beirle et al., 2004; Schaub et al., 2007).

This estimated  $NO_x$  lifetime by the column model is to a large extent determined by the decline in the line density downwind of the source area. Here four effects were tested that influence the downwind decay that were not yet incorporated in the column model: downwind emissions, a diurnal emission cycle, variability in lifetime and background decay. Downwind emissions that continuously add  $NO_2$  caused a slower decline in the line density and hence a longer estimated lifetime. Accounting for this effect resulted in the longest lifetime from the four experiments: 3 times longer than the standard column model, resulting in a  $NO_x$  emission that is only 7.1% lower than the reference case. Hence it can be concluded that these four effects have large impacts on the estimated  $NO_x$  lifetime, but that their impact on the estimated  $NO_x$  emission is only minor. This can be explained by the fact that the emission is defined by the increase in line density over the city and that the emission is dominant compared to chemical decay. Yet it remains difficult to obtain an accurate estimate of the  $NO_x$  lifetime by fitting an exponential function through the downwind decay in the line density.

Because of this complexity in obtaining the  $NO_x$  lifetime from the downwind decay, the  $NO_x$  emission was estimated just from the accumulation over the city with a prescribed lifetime from a chemical reanalysis. This was done not only for 22 November 2017, but also for four days in February and six days in April 2018. Surface observations in Paris for these days showed a similar variation in magnitude in space as the TROPOMI observations, which confirms the increase of  $NO_2$  in the wind direction in the line densities. The resulting averaged  $NO_x$  emission estimates are twice as strong during weekdays (55.9 mol s<sup>-1</sup>) than during weekends (30.2 mol s<sup>-1</sup>) and respectively 24% and 51% lower than the EDGAR emission inventory for Paris from 2012 with temporal correction. Although these lower values can partly be explained by the reported  $NO_x$  reduction since 2012, it should be noted that comparing these top-down estimates with the bottom-up inventory requires caution, as the uncertainty in the temporal correction factors seems to be large.

It is recommended to further improve this research by validating the TROPOMI observations, analysing the vertical distribution of  $NO_x$  over a source area such as Paris and how this relates to the wind speed, and replacing the average  $NO_x/NO_2$  ratio with either simulated or measured values. But most importantly, extending this approach for more days and to other source areas could give better insights into the variation of the  $NO_x$  emission in time and around the world.

Once this method is applied more widely and further evaluated, it could become an understandable and fast addition to the existing bottom-up emission accounting and provide better insight into the magnitude, as well as the spatial and temporal variation of  $NO_x$  emissions. An important advantage is that this method is easy to understand: the TROPOMI observations give a good visualisation of how  $NO_x$  is added to the atmosphere from a large emission source. Yet

the strongest point of this approach is  $NO_x$  emissions can be quantified day-to-day, given that the source area is not covered by clouds and the wind speed is not very low. This allows to examine the seasonal and weekly emission cycle, from which the contribution of different emission sources can be deduced since emissions from residential heating take place during the winter and traffic emissions are strongest during weekdays. This can be an important tool for evaluating air quality and climate mitigation measures, not only for Paris' new atmospheric protection plan, but also for other large  $NO_x$  emissions sources all around the world.

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